

# Evolution in magma storage conditions beneath Kick-'em-Jenny and Kick-'em-Jack submarine volcanoes, Lesser Antilles arc

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## ABSTRACT

Kick-'em-Jenny and Kick-'em-Jack are two submarine volcanoes located in the southern Lesser Antilles arc. We present data on the temporal and spatial compositional variations in lavas and melt inclusions from Kick-'em-Jenny and four lavas from Kick-'em-Jack. Analyses of crater samples from Kick-'em-Jenny show an evolution in magma compositions between 1972 and 2013/14. Rock types have progressed from predominantly basalts in 1972 to basaltic andesites in 2013/14, while melt inclusions have also become more silicic over time, ranging basaltic to andesitic in 1972 and basaltic andesite to dacitic in 2013/14. Hornblende-phyric basalts are unusually rich in MgO ( $\leq 13$  wt%) and Ni ( $\leq 230$  ppm) and contain rare olivine phenocryst cores of Fo<sub>90–92</sub>, suggestive of a primitive character. Other phenocryst phases include spinel, clinopyroxene and plagioclase. Conversely, 2013/14 landslide deposit samples contain orthopyroxene as an additional phenocryst phase and are distinct from crater deposits in terms of phenocryst and bulk rock chemistry. We propose that 1972 and 2013/14 crater deposits sample Kick-'em-Jenny's current edifice, whereas 2013/14 landslide deposits likely represent "proto Kick-'em-Jenny" (an edifice existing pre 43.5 kyr BP). Geochemical data for Kick-'em-Jack support the notion that it is part of the same volcanic complex as Kick-'em-Jenny and probably represents part of the now collapsed proto Kick-'em-Jenny.

Low eruption temperatures ( $\sim 1000$  °C), estimated from geothermometry, coupled with whole-rock magma chemistry suggest an origin of the most mafic lavas of Kick-'em-Jenny by dehydration melting of hornblende-peridotite just below the Moho. Subsequent injection of these cool, oxidised, wet primitive basalts into the deep crust was accompanied early fractionation of hornblende-dominated assemblages. Large, zoned hornblende phenocrysts track rapid ascent of magmas through the crust along a trajectory sub-parallel to the hornblende liquidus curve. Melt inclusion data provide clear evidence of volatile-rich ( $\leq 5$  wt% H<sub>2</sub>O and  $< 3200$  ppm CO<sub>2</sub>) pre-eruptive magma storage 7–12 km below Kick-'em-Jenny. This new understanding of the architecture of Kick-'em-Jenny's sub-volcanic system will be useful for interpreting future seismic signals and mitigating volcanic hazards.

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## 1. Introduction

Underwater eruptions are the most common type on Earth, yet little is known about submarine volcanoes compared to their subaerial counterparts. This is largely due to the difficulty in making direct observations of underwater eruptions. Monitoring of submarine volcanoes is typically limited to land-based seismic networks and reconnaissance ship surveys. A key drawback to our understanding of both submarine and subaerial volcanoes is our inability to directly observe the

configuration of magma reservoirs and the processes operating beneath volcanoes. Technological advancements have made the imaging of sub-volcanic systems possible through geophysical surveys, however the acquisition of such measurements is often contingent on access to equipment and resources that are not often readily available. Petrological and geochemical observations of volcanic rock samples provide an alternate means of retrospective imaging of crustal reservoirs as crystals and their entrapped melts record the changing chemical and thermal conditions occurring within reservoirs before eruption (Cooper, 2017).

Here, we focus on lava samples from two submarine volcanoes in the Lesser Antilles arc: Kick-'em-Jenny and Kick-'em-Jack. We integrate new petrological and geochemical data for lavas and phenocryst-hosted melt inclusions with seismic data to constrain the vertical extent

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of sub-volcanic magmatic systems and pre-eruptive magma storage conditions with implications for monitoring and hazard assessment.

### 1.1. Volcanic activity and previous research

Kick-'em-Jenny and Kick-'em-Jack mark the southernmost culmination of volcanism in the Grenadines archipelago of the Lesser Antilles arc (Fig. 1), a 750 km-long, intra-oceanic island chain resulting from the slow westward subduction (~2 cm/yr) of the North and South American plates beneath the Caribbean plate. Magma production rates for this arc are comparatively low (Wadge, 1984), with only nine of twenty one potentially active volcanoes displaying historical eruptive activity (Camejo and Robertson, 2013).

#### 1.1.1. Kick-'em-Jenny

Located 8 km north of Grenada, Kick-'em-Jenny is the only known active submarine volcano in the Lesser Antilles (Lindsay and Shepherd, 2005). Its asymmetric cone, 197 m below sea level, is the result of irregular periods of constructive edifice building during eruptive episodes interspersed with mass wasting (Allen et al., 2018). Although the volcano's existence first became known following felt explosions related to an eruption in 1939 (Devas and MacAdam-Sherwin, 1939), monitoring by the University of the West Indies, Seismic Research Centre (UWI-SRC) only began when this agency was established in 1953. Fourteen eruptions have been documented (Table 1), detected by human observation and/or seismometer recordings of T-phase episodes (Lindsay and Shepherd, 2005). The most recent detected eruption was in 2017, making it the most historically active volcano of the entire arc.

Kick-'em-Jenny's proclivity for edifice instability is evident from its morphology and landslide deposits. The current cone lies within a

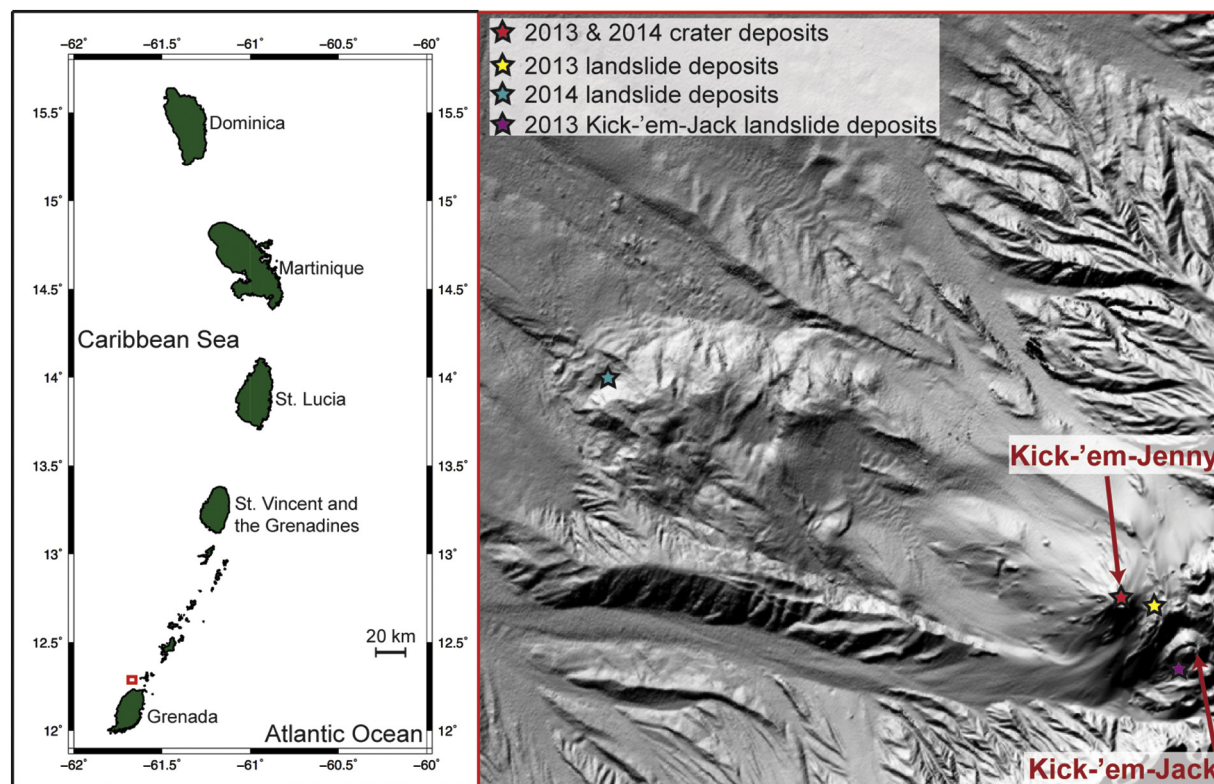
**Table 1**

Historical eruptions at Kick-'em-Jenny volcano.  
(Modified after Lindsay and Shepherd (2005).)

| Eruption date             | Evidence for eruption |                  |                                                                     |
|---------------------------|-----------------------|------------------|---------------------------------------------------------------------|
|                           | T-phase signal        | Felt earthquakes | Surface manifestations: water disturbances, ejected material, gases |
| July 24, 1939             | –                     | ✓                | ✓                                                                   |
| Oct. 5–6, 1943            | ✓                     | ✓                | –                                                                   |
| Oct. 30, 1953             | ✓                     | ✓                | –                                                                   |
| Oct. 24, 1965             | ✓                     | ✓                | –                                                                   |
| May 5–7, 1966             | ✓                     | –                | –                                                                   |
| Aug. 3–6, 1966            | ✓                     | ✓                | –                                                                   |
| July 5, 1972              | ✓                     | –                | –                                                                   |
| Sept. 6, 1974             | ✓                     | –                | ✓                                                                   |
| Jan. 14, 1977             | ✓                     | –                | –                                                                   |
| Dec. 29–30, 1988          | –                     | ✓                | ✓                                                                   |
| March 26 to April 5, 1990 | ✓                     | ✓                | –                                                                   |
| Dec. 4–6, 2001            | ✓                     | ✓                | –                                                                   |
| July 23–24, 2015          | ✓                     | ✓                | –                                                                   |
| April 29, 2017            | ✓                     | ✓                | –                                                                   |

Data taken from Latchman et al. (2017), Robertson et al. (2015) and Lindsay and Shepherd (2005).

horseshoe-shaped collapse scar (formed at least 43.5 kyr BP based on radiocarbon dating, Dondin et al. (2017)), which is open to the west (Lindsay et al., 2005), with numerous associated landslide deposit sequences downslope (Fig. 1). This horseshoe feature is thought to represent the remains of the near-total collapse of an ancestral “proto Kick-'em-Jenny” edifice from which the current edifice now emerges. The proto edifice is modelled to have stood above sea-level before its subsequent failure (Dondin et al., 2012). Recent evaluation of bathymetric data indicates that over the past thirty years, cone-building episodes



**Fig. 1.** Map showing location of Kick-'em-Jenny and Kick-'em-Jack volcanoes. Bathymetry is derived from multibeam data collected on the E/V Nautilus NA039 cruise in 2013. Inset map shows location of these volcanoes (red rectangle) in relation to the southern half of the Lesser Antilles arc. Stars denote locations of samples collected from Kick-'em-Jenny and Kick-'em-Jack in 2013 and 2014. 1972 samples (not shown) are reported to have been collected from various parts of the cone and crater (Sigurdsson and Shepherd, 1974).

have been infrequent with significantly more material being shed from the edifice during debris avalanche events triggered by seismicity or explosive eruptions (Allen et al., 2018). Although submarine landslides can be a key contributor to tsunami generation, the scientific consensus remains that the volcano's current vent height is insufficient to generate a tsunami (Allen et al., 2018; Lindsay et al., 2005). Of more immediate concern are the direct hazards associated with the volcano, such as ballistic ejecta, water disturbances and lowered water-column density due to degassing (Lindsay et al., 2005).

Kick-'em-Jenny erupts both explosively and effusively. The deciding mechanism for which style of activity will characterize a given eruption remains unclear. Estimated H<sub>2</sub>O contents of pre-eruptive magmas at Kick-'em-Jenny suggest that they are enriched in volatiles while stored at magma reservoir depths. Devine and Sigurdsson (1995) alluded to the amount of pre-eruptive degassing of volatiles from saturated magma determining the style of activity: efficient magma degassing produces effusive eruptions while significant retention of dissolved volatiles generates explosive eruptions. The extent to which magma can degas depends ultimately on the rate of magma ascent and the permeability of the plumbing system. Currently, the former mechanism remains poorly understood, but the latter has been tackled by deformation experiments on rock samples collected from Kick-'em-Jenny. Fractures created by brittle deformation of edifice rocks promote efficient outgassing, favouring lava dome formation on this volcano (Dondin et al., 2017). Gas escape through interconnected bubble networks within magmas can also produce extrusive activity (Melnik et al., 2005). However, triggering of explosive behaviour by rapid magma ascent, irrespective of edifice permeability (Dondin et al., 2017), as well as magma interaction with external water (Devine and Sigurdsson, 1995) cannot be ruled out.

Petrological studies on Kick-'em-Jenny have been scarce due to the availability of rock samples. The earliest petrological studies were of rocks collected on the Royal Navy oceanographic research vessel *H.M.S. Hecla* in May 1972 (Devine and Sigurdsson, 1995; Sigurdsson and Shepherd, 1974). The principal rock types are scoriaceous, porphyritic basalt and minor basaltic andesites and andesites with abundant hornblende phenocrysts. There are few examples globally of hornblende bearing basalts e.g. Rhon, Germany (Mayer et al., 2013); Cerro La Pilita, Western Mexico (Luhr and Carmichael, 1985); Mount Lamington, Papua New Guinea (Arculus et al., 1983); Bogoslof, Alaska (Arculus et al., 1977). The presence of hornblende as a basaltic phenocryst has prompted interest in the water content of primary magmas in island arcs and the role of hornblende in crustal differentiation (Devine and Sigurdsson, 1995). Volcanic sediment collected on a 2003 *R/V Ronald H. Brown* cruise from areas of hydrothermal activity has also been studied in the context of hydrothermal alteration (Olsen, 2011).

#### 1.1.2. Kick-'em-Jack

Kick-'em-Jack, located 3 km east of Kick-'em-Jenny (Fig. 1), was first discovered during the scientific cruise by the *R/V Ronald H. Brown* in 2003 (Carey et al., 2014; Lindsay and Shepherd, 2005). Bathymetric mapping revealed that Kick-'em-Jack, along with four other submarine craters and domes, lies towards the east of Kick-'em-Jenny. These features are inferred to be older than Kick-'em-Jenny and no longer active (Lindsay and Shepherd, 2005), however some low-level hydrothermal activity was observed between Kick-'em-Jenny and Kick-'em-Jack in 2003 (Carey et al., 2014). Kick-'em-Jack is described as similar in size to Kick-'em-Jenny with its own horseshoe-shaped crater enclosing a dome (Carey et al., 2014). The close association of Kick-'em-Jack with Kick-'em-Jenny suggests that they may both belong to the same volcanic complex including the now-collapsed proto Kick-'em-Jenny volcano (Lindsay and Shepherd, 2005). The first and only petrological description to date of deposits from Kick-'em-Jack was provided by Christopher et al. (2016) on rock samples collected on the NA039 research cruise on the *E/V Nautilus* in 2013. Samples analysed were basaltic andesites and andesites containing up to 35 vol% phenocrysts of plagioclase, hornblende and minor clinopyroxene.

## 2. Methods

### 2.1. Sampling

For this study, lava samples were obtained from the Exploration Vessel *E/V Nautilus* cruises NA039 during 1–18 November 2013 and NA054 during September 21 to October 9, 2014 using a remotely operated vehicle. From the 2013 cruise, we analysed eleven samples from Kick-'em-Jenny's crater and three from a nearby debris avalanche deposit (Fig. 1). A further four samples were analysed from debris avalanche deposits associated with Kick-'em-Jack. These represent a subset of the first documented collection of samples from Kick-'em-Jack volcano. From the 2014 cruise, we analysed four samples collected from Kick-'em-Jenny's crater and five from a landslide deposit (Fig. 1). We infer that debris avalanche/landslide deposits represent collapsed material from one or more earlier eruptions of the respective volcanoes. We re-analysed five samples collected using grab sampling by the *H.M.S. Hecla* survey in 1972, first described in Sigurdsson and Shepherd (1974), and kindly provided by Steve Carey (University of Rhode Island). These samples were reportedly taken from various parts of the cone and crater (Sigurdsson and Shepherd, 1974). Sample collections from 1972 and 2013/14 expeditions occurred ~6 and ~12 years, respectively, after the corresponding documented eruption (Table 1). There is some duplication in names between 1972 and 2013/14 Kick-'em-Jenny samples (e.g. KEJ010, Fig. 2) used in this study. Where this occurs, the date of collection is stipulated in parentheses, e.g. KEJ010 (1972) or KEJ010 (2014).

### 2.2. Whole rock chemistry

Whole rock major and trace elements were analysed by X-ray fluorescence spectrometry (XRF) on twenty Kick-'em-Jenny and three Kick-'em-Jack samples using a Bruker-AXS S4 Pioneer instrument at the University of East Anglia. Samples were first crushed and powdered using a rock splitter, jaw crusher and planetary ball mill at the University of Bristol. Pressed pellets and glass beads were then prepared at the University of East Anglia for analysis. For major elements, the instrument was calibrated using the following standards: SDO-1, GSP-2, W2a, AC-E and AGV-2. For trace elements, the instrument was calibrated using the following standards: SDO-1, Mess-2, STSD-2 and W2a. Differences between measured and reference values for these standards are <0.44 wt% and <8 ppm for major and trace elements respectively.

### 2.3. Mineral and melt inclusion chemistry

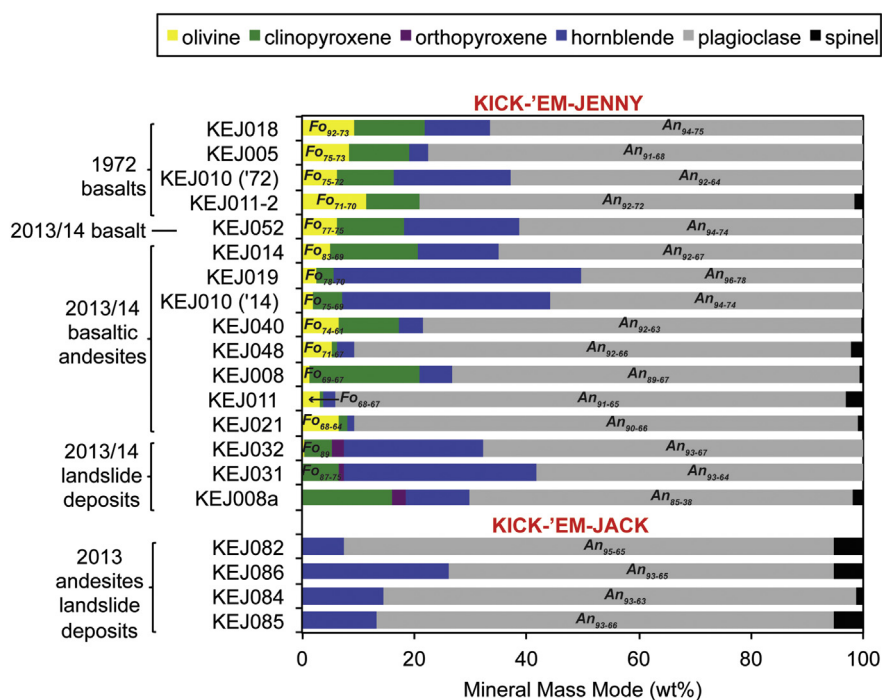
#### 2.3.1. Major elements

Minerals and melt inclusions (MIs) were analysed using a Cameca SX100 electron microprobe at the University of Bristol. Samples were analysed as polished, carbon-coated thin sections. Some minerals were also hand picked from coarsely crushed samples, mounted in epoxy, polished until MIs were exposed and then carbon-coated for analysis. For mineral analyses, a 20 kV accelerating voltage, 10 nA beam current and 1 µm spot were used. For MI analyses, a 20 kV accelerating voltage, 4 nA beam current and 5 µm-defocused spot were used to minimize alkali migration during analysis (Humphreys et al., 2006). Both techniques were calibrated using oxide, mineral and metal standards.

#### 2.3.2. Trace elements and volatiles

Trace element, H<sub>2</sub>O and CO<sub>2</sub> concentrations in MIs were analysed on gold-coated, polished thin sections and grain mounts by secondary ion mass spectrometry (SIMS) using the Cameca IMS4f ion-microprobe at the NERC facility, University of Edinburgh. Carbon coats were thoroughly removed prior to SIMS analysis. On any one MI, CO<sub>2</sub> (as <sup>12</sup>C<sup>+</sup>) was measured first, followed by H<sub>2</sub>O (as <sup>1</sup>H<sup>+</sup>) and trace elements (<sup>7</sup>Li, <sup>11</sup>B, <sup>47</sup>Ti, <sup>88</sup>Sr, <sup>89</sup>Y, <sup>90</sup>Zr, <sup>93</sup>Nb, <sup>138</sup>Ba, <sup>140</sup>Ce) in a separate routine. Both





**Fig. 2.** Modal proportions by mass of Kick-'em-Jenny and Kick-'em-Jack phenocrysts. Samples are listed in order of decreasing Fo content of olivine followed by An content of plagioclase within a specific rock type group. 2013/14 landslide deposits consist of two basaltic andesites and one andesite (KEJ008a).

analytical routines were conducted on exactly the same area of the inclusion. Instrument conditions comprised 10 kV  $^{16}\text{O}^-$  2–5 nA primary beam focused to an  $\sim 10\ \mu\text{m}$  spot on the sample surface. Positive secondary ions were extracted at 4.5 kV with an energy offset of 50 eV for  $\text{CO}_2$  (75 eV for  $\text{H}_2\text{O}$ ) and a  $25\ \mu\text{m}$  image field.  $^{12}\text{C}^+$  data were obtained at high mass resolution to eliminate  $^{24}\text{Mg}^{2+}$  interference. Volatile species were calibrated against basaltic glass standards using data analysis procedures of Blundy and Cashman (2008). Trace elements were calibrated against glass standard GSD.

### 3. Results

#### 3.1. Petrography

Kick-'em-Jenny 1972 samples consist of basalts, while the 2013/14 collection consists of basalts, basaltic andesites and andesites (Fig. 2). It should be noted that non-vesicular basaltic andesites and andesites were also collected during the 1972 expedition (Devine and Sigurdsson, 1995), but were proportionally less abundant than basalts (Supplementary Fig. S1). All rock types are glomeroporphyritic and vesicular (up to 16% by volume) with phenocrysts of olivine, clinopyroxene, orthopyroxene, hornblende, plagioclase and spinel (Fig. 3a–f). Some minerals show embayed or corroded edges as well as resorption of grain interiors. Mineral textures are summarised in Table 2. Plagioclase is the most abundant phase. Hornblende and clinopyroxene are the second most abundant phases in 56% and 44% of samples respectively. Orthopyroxene is restricted to samples collected from debris avalanche deposits, while being absent from crater-derived samples. Crystal clusters, showing a variety of grain size arrangements, are ubiquitous (Fig. 3c & d). Polyminerals containing two or more phases are typical, with fewer instances of monomineralic (clinopyroxene- and plagioclase-only) clots. Recrystallised interstitial glass is observed in some clusters (Fig. 3c). Some constituent minerals (e.g. hornblende) also appear to be breaking down when in contact with glass or other minerals. Kick-'em-Jenny's lavas quench upon eruption under water preventing extensive degassing and readily preserving melt inclusions in phenocryst phases.

Secondary minerals such as calcite and prehnite occupy vesicles and/or veins in some samples (Fig. 3c), suggesting that some Kick-'em-Jenny lavas were affected by low temperature burial metamorphism and seawater infiltration. The groundmass consists of microlites of spinel, olivine, plagioclase, clinopyroxene, orthopyroxene and apatite together with some glass, collectively making up some 38–70% by volume on a vesicle-free basis. The groundmass of one sample (KEJ010–2014) displays an irregular interplay of predominantly dark and light minerals suggesting compositional variability of the melt (Fig. 3f) and possibly mixing of magmas and/or entrained crystals.

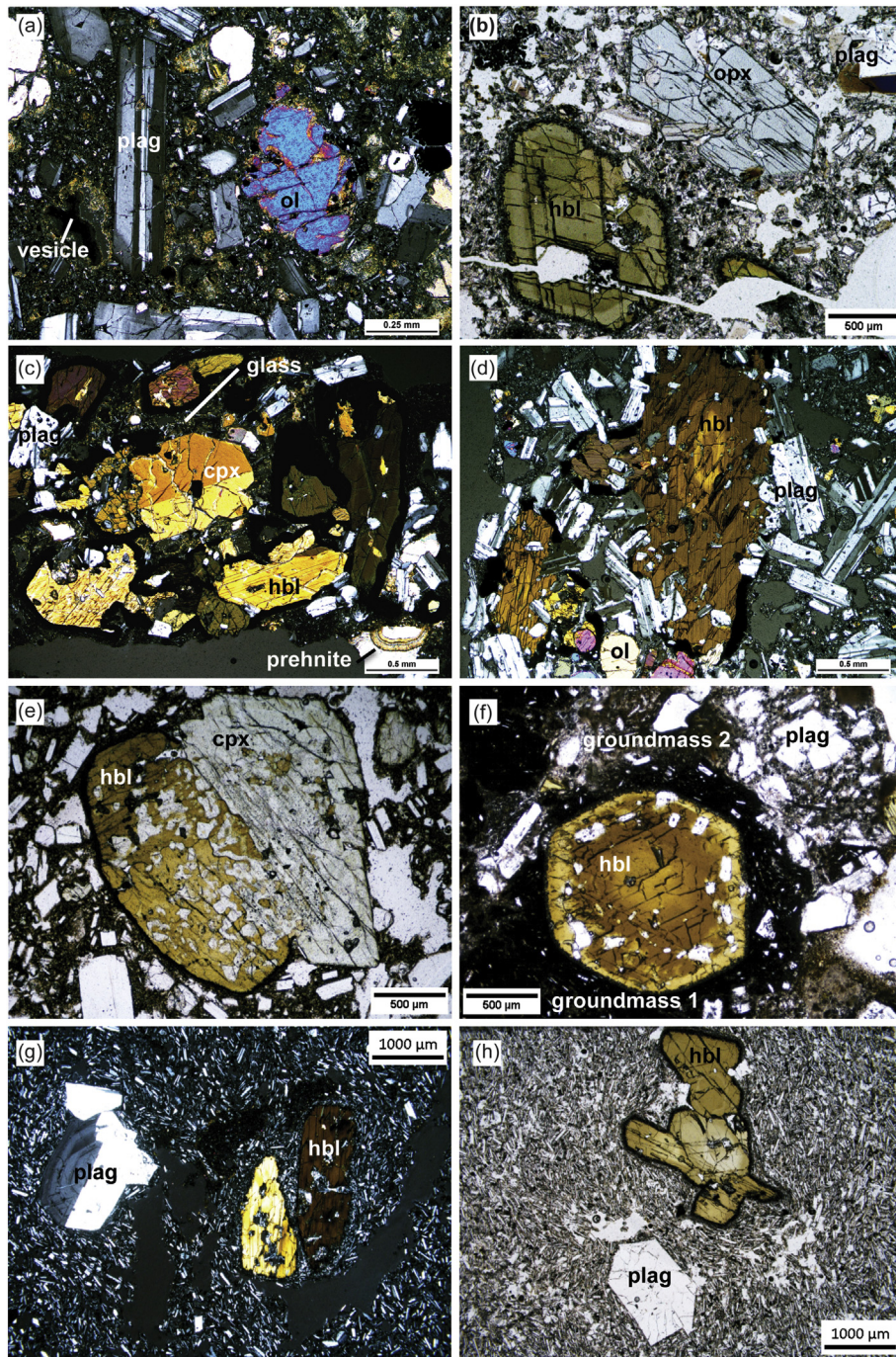
Kick-'em-Jack andesites are porphyritic, consisting of mainly hornblende, plagioclase and spinel phenocrysts (Fig. 3g & h, Table 2). Olivine was observed in one sample included in a hornblende grain. Again, plagioclase is the predominant phase, followed by hornblende (Fig. 2). Glomerocrysts are less prevalent in Kick-'em-Jack lavas than in Kick-'em-Jenny lavas, but when present exhibit both poly- and monomineralic clusters (Fig. 3h). Vesicles make up  $\leq 6\%$  by volume. The groundmass consists of microlites of plagioclase and spinel embedded in glass making up 89–93% by volume on a vesicle-free basis.

#### 3.2. Mineral chemistry

Full mineral composition datasets for Kick-'em-Jenny and Kick-'em-Jack lavas are presented in Supplementary Tables S1 and S2 respectively. Ferric iron contents were estimated using the stoichiometric methods of Droop (1987) for spinel, Lindsley (1983) for clinopyroxene, and Holland and Blundy (1994) for hornblende. For spinel and clinopyroxene, Mg# is expressed as  $100\ \text{Mg}/(\text{Mg} + \text{Fe}^{2+})$  and for hornblende, Mg# is expressed as  $100\ \text{Mg}/(\text{Mg} + \text{Fe}^{\text{Total}})$ .

##### 3.2.1. Olivine

Kick-'em-Jenny olivine phenocryst compositions range from  $\text{Fo}_{92-61}$  (Fig. 2).  $\text{Fo}_{<70}$  olivines are restricted to 2013/14 samples, while select 1972 (KEJ018), 2013 (KEJ014) and 2013/14 landslide deposit (KEJ031 and KEJ032) samples contain the highest Fo contents ( $>80$ ) (Fig. 5). Interestingly, these samples also exhibit significant normal zoning, e.g. KEJ018 (e.g.  $\text{Fo}_{92-79}$ ), KEJ014 (e.g.  $\text{Fo}_{83-75}$ ) and KEJ032 (e.g.  $\text{Fo}_{87-80}$ )



**Fig. 3.** Photomicrographs of representative minerals and textures in Kick-em-Jenny (a–f) and Kick-em-Jack lavas (g–h) (cross polarised light, xpl; plane polarised light, ppl). (a) Basaltic andesite with olivine altered along grain edges and fractures, sieve textured plagioclase and vesicles, xpl (KEJ014). (b) Basaltic andesite with orthopyroxene and zoned hornblende, xpl (KEJ031). (c) Crystal cluster with recrystallised interstitial melt, xpl (KEJ049). (d) Cluster with millimeter and micrometer sized crystals, xpl (KEJ008). (e) Crystal clots displaying varying degrees of uraltization: left crystal shows clinopyroxene mostly replaced to hornblende and right crystal showing the beginnings of clinopyroxene replacement (KEJ018), ppl. (f) Groundmass compositional variability, ppl (KEJ010, 2014). (g) Typical oscillatory-zoned plagioclase with trachytic groundmass, xpl (KEJ084). (h) Hornblende-dominated glomerocryst, ppl (KEJ086).

with the most Fo-rich compositions preserved as small, corroded grain cores with abrupt transitions to less Fo-rich rims (Fig. 4b & d). Olivine grains elsewhere (phenocrysts and glomerocrysts) are only marginally zoned (normal and reverse) showing at most a 1% Fo rimward decrease or increase (Fig. 4a). There is a steady decrease in MnO from 0.8 to 0.15 wt% with decreasing Fo (Supplementary Table S1a). NiO (0.03 to 0.50 wt%) increases with increasing Fo (Fig. 5a). Landslide deposits are distinct in their trend with higher NiO contents for a given MgO. CaO varies between 0.11 and 0.30 wt% with no discernible relationship with Fo (Fig. 5b). CaO contents are always higher in phenocryst rims

compared to cores showing for the most part a rimward increase of 0.06 wt% for 1972 and 0.03 wt% for 2013/14 lavas. Phenocrysts and glomerocrysts follow a similar distribution with no significant compositional distinction. Microlite compositions range Fo<sub>66–59</sub>, <0.02 wt% NiO and 0.34–0.41 wt% CaO (Fig. 5). Also plotted on Fig. 5b are isotherms calculated using the Ca-in-olivine thermometer of Shejwalkar and Coogan (2013) appropriate for orthopyroxene-bearing lavas KEJ031 and KEJ032.

The single olivine grain analysed from Kick-em-Jack has Fo<sub>79</sub>, CaO 0.09 wt% and NiO 0.13 wt%.



**Table 2**  
Phenocryst petrographic descriptions for Kick-'em-Jenny and Kick-'em-Jack lavas.

| Mineral phase           | Volcano                        | Grain size & shape                                                                                                                     | Textural types                                                                                                                                                                                                                                                                                                                                                                                                                                            | Other comments                                                                                                                                                                                                                                                        |
|-------------------------|--------------------------------|----------------------------------------------------------------------------------------------------------------------------------------|-----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|-----------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| Olivine                 | Kick-'em-Jenny                 | <ul style="list-style-type: none"> <li>&lt;500 <math>\mu\text{m}</math>, KEJ014 &lt; 1 mm</li> <li>Subhedral &amp; anhedral</li> </ul> | <ul style="list-style-type: none"> <li>Isolated phenocrysts and microphenocrysts (Fig. 3a)</li> <li>Clotted with spinel, clinopyroxene, orthopyroxene, hornblende and/or plagioclase (Fig. 3d)</li> </ul>                                                                                                                                                                                                                                                 | <ul style="list-style-type: none"> <li>Zoned</li> <li>Unaltered except KEJ014 (Fig. 4d)</li> <li>Rare in debris avalanche deposits (&lt;1%) forming clots with and surrounded by overgrowths of clinopyroxene, hornblende (Fig. 4c) and orthopyroxene</li> </ul>      |
| Clinopyroxene           | Kick-'em-Jenny                 | <ul style="list-style-type: none"> <li>&lt;2.5 mm</li> <li>Euhedral, subhedral &amp; anhedral</li> </ul>                               | <ul style="list-style-type: none"> <li>Large isolated euhedral/subhedral grains (~2.5 mm)</li> <li>Smaller isolated euhedral/subhedral grains (&lt;2 mm)</li> <li>Subhedral/anhedral grains (&lt;2 mm) clotted with olivine, hornblende, plagioclase and/or spinel</li> <li>Grains showing replacement to hornblende (Fig. 3e)</li> </ul>                                                                                                                 | <ul style="list-style-type: none"> <li>Zoned (compositional, oscillatory and sector)</li> <li>Inclusions of spinel and plagioclase</li> </ul>                                                                                                                         |
| Orthopyroxene           | Kick-'em-Jenny                 | <ul style="list-style-type: none"> <li>&lt;2 mm</li> <li>Subhedral</li> </ul>                                                          | <ul style="list-style-type: none"> <li>Isolated grains (Fig. 3b)</li> <li>Clotted with olivine, clinopyroxene and/or plagioclase</li> </ul>                                                                                                                                                                                                                                                                                                               | <ul style="list-style-type: none"> <li>Zoned</li> <li>Present only debris avalanche deposits</li> </ul>                                                                                                                                                               |
| Hornblende              | Kick-'em-Jenny                 | <ul style="list-style-type: none"> <li>&lt;5 mm</li> <li>Euhedral, subhedral &amp; anhedral</li> </ul>                                 | <ul style="list-style-type: none"> <li>Euhedral/subhedral isolated phenocrysts with or without inclusions of olivine, clinopyroxene, plagioclase (Fig. 3f)</li> <li>Anhedral/subhedral isolated phenocrysts with fragmented/resorbed grain interiors and edges (Fig. 3b)</li> <li>Clotted with plagioclase, clinopyroxene and/or orthopyroxene (Fig. 3c &amp; d)</li> <li>Grains showing uraltization of clinopyroxene to hornblende (Fig. 3e)</li> </ul> | <ul style="list-style-type: none"> <li>Zoned phenocrysts (compositional and oscillatory), unzoned microphenocrysts</li> <li>Absent from KEJ011–2 (1972)</li> <li>Thick opacite reaction rims (Fig. 3b, c, d &amp; e), few grains are completely opacitized</li> </ul> |
|                         | Kick-'em-Jack                  | <ul style="list-style-type: none"> <li>&lt;5 mm</li> <li>Euhedral, subhedral &amp; anhedral</li> </ul>                                 | <ul style="list-style-type: none"> <li>Large (~5 mm), isolated, euhedral/subhedral phenocrysts with/without inclusions of other phases</li> <li>Smaller isolated euhedral/subhedral phenocrysts (&lt;2 mm) with/without inclusions of other phases (Fig. 3g)</li> <li>Anhedral/subhedral phenocrysts clotted with other hornblende grains and/or other phases (Fig. 3h)</li> </ul>                                                                        | <ul style="list-style-type: none"> <li>Zoned phenocrysts, unzoned microphenocrysts</li> <li>Opacite rims (Fig. 3h)</li> </ul>                                                                                                                                         |
| Plagioclase             | Kick-'em-Jenny                 | <ul style="list-style-type: none"> <li>&lt;2.5 mm</li> <li>Subhedral</li> </ul>                                                        | <ul style="list-style-type: none"> <li>Large isolated phenocrysts (~2–2.5 mm) with thinly zoned rims</li> <li>Isolated phenocrysts &amp; microphenocrysts (&lt;2 mm) (Fig. 3a)</li> <li>Clotted with other minerals (Fig. 3c &amp; d)</li> </ul>                                                                                                                                                                                                          | <ul style="list-style-type: none"> <li>Zoned (oscillatory only in isolated phenocrysts)</li> <li>Sieve textured cores (except some microphenocrysts)</li> <li>Minor spinel, olivine and hornblende inclusions</li> </ul>                                              |
|                         | Kick-'em-Jack                  | <ul style="list-style-type: none"> <li>&lt;2 mm</li> <li>Subhedral</li> </ul>                                                          | <ul style="list-style-type: none"> <li>Isolated phenocrysts/micropheocrysts</li> <li>Clotted with other plagioclase grains, hornblende and/or spinel</li> </ul>                                                                                                                                                                                                                                                                                           | <ul style="list-style-type: none"> <li>Zoned (oscillatory) (Fig. 3g)</li> <li>Sieve textured cores</li> </ul>                                                                                                                                                         |
| Spinel (magnetite-rich) | Kick-'em-Jenny & Kick-'em-Jack | <ul style="list-style-type: none"> <li>&lt;200 <math>\mu\text{m}</math></li> <li>Subhedral &amp; anhedral</li> </ul>                   | <ul style="list-style-type: none"> <li>Isolated</li> <li>Clotted with other phases</li> </ul>                                                                                                                                                                                                                                                                                                                                                             | <ul style="list-style-type: none"> <li>Included in all phases</li> </ul>                                                                                                                                                                                              |

### 3.2.2. Pyroxene

All textural types of phenocryst clinopyroxene (Table 2) in Kick-'em-Jenny lavas have diopside and high-Ca augite compositions with  $\text{Mg}\#_{93-71}$  (Supplementary Table S1c). Zoning is ubiquitous with analyses from both time periods displaying normal (maximum 15 mol% rimward decrease, e.g. KEJ018  $\text{Mg}\#_{88-73}$ ) and reverse (maximum 6 mol% rimward increase, e.g. KEJ005,  $\text{Mg}\#_{79-85}$ ) zoning. Ca content in clinopyroxene increases (0.73–0.93 p.f.u.) with decreasing  $\text{Mg}\#$ , but there is no clear chemical distinction between phenocrysts and glomerocrysts for either 1972 or 2013/14 deposits.  $\text{Al}^{\text{IV}}$  contents vary between 0.04 and 0.24 p.f.u. and  $\text{TiO}_2$  varies between 0.2 and 1.1 wt%. There is no clear correlation between  $\text{Mg}\#$  and  $\text{Al}^{\text{IV}}$  or  $\text{TiO}_2$ . Microlite compositions are also diopside and high-Ca augite, mirroring trends of phenocryst Ca (0.60–0.85 p.f.u.),  $\text{Al}^{\text{IV}}$  (0.01–0.19 p.f.u.) and  $\text{TiO}_2$  (0.2–1.5 wt%) with  $\text{Mg}\#$  ( $\text{Mg}\#_{100-63}$ ).

Phenocryst and microlite orthopyroxene for Kick-'em-Jenny lavas are enstatite in composition ( $\text{Mg}\#_{78-67}$  and  $\text{Mg}\#_{75-68}$  respectively) (Supplementary Table S1d). Normally zoned phenocrysts display up to a 5 mol% rimward decrease in  $\text{Mg}\#$  (e.g.  $\text{Mg}\#_{77-72}$ , KEJ031). Phenocryst  $\text{Al}^{\text{IV}}$  compositions (0.01–0.11) increase with increasing  $\text{Mg}\#$ , while Ca (p.f.u.) and Ti (p.f.u.) compositions show no discernible pattern with  $\text{Mg}\#$ , with contents of <0.10 and <0.16 respectively. Microlite  $\text{Al}^{\text{IV}}$  compositions (<0.03) also increase with increasing  $\text{Mg}\#$ , while Ca (<0.06 p.f.u.) and  $\text{TiO}_2$  (<0.3 wt%) compositions show no discernible pattern with  $\text{Mg}\#$ .

### 3.2.3. Hornblende

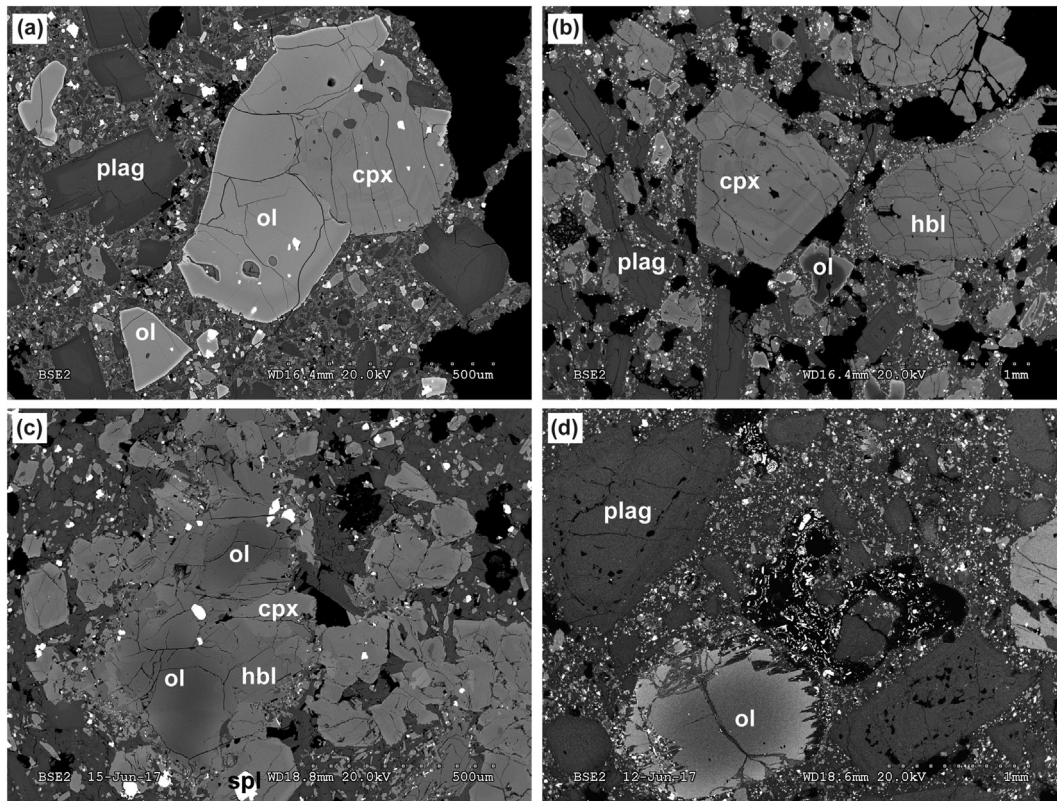
Phenocryst hornblendes in Kick-'em-Jenny and Kick-'em-Jack lavas are pargasite according to Leake et al. (2003) with  $\text{Mg}\#$  ranges of 80–62 and 79–63 respectively (Supplementary Tables S1e and S2c).

For Kick-'em-Jenny, normally zoned phenocrysts show up to a 9%  $\text{Mg}\#$  rimward decrease, e.g.  $\text{Mg}\#_{89-80}$ , KEJ008 (up to 4% for Kick-'em-Jack glomerocrysts, e.g.  $\text{Mg}\#_{89-85}$ , KEJ085), while reversely zoned grains show up to a 11%  $\text{Mg}\#$  rimward increase, e.g.  $\text{Mg}\#_{83-94}$ , KEJ032 (up to 9% for Kick-'em-Jack glomerocrysts, e.g.  $\text{Mg}\#_{84-93}$ ). A striking feature of Kick-'em-Jenny and Kick-'em-Jack hornblendes are their high  $\text{Al}_2\text{O}_3$  contents (10.7–15.6 wt%). Samples from both volcanoes show a decrease in K and  $\text{Al}^{\text{IV}}$  with decreasing  $\text{Mg}\#$  (Fig. 6a & c). Most samples display a fairly constant Ti concentration with changing  $\text{Mg}\#$ , while a subset of 1972 and 2013/14 edifice deposits display a contrasting trend of decreasing Ti content with increasing  $\text{Mg}\#$  (Fig. 6b). 1972 and 2013/14 edifice deposit hornblendes have distinct compositions in terms of K, Ti and  $\text{Al}^{\text{IV}}$ , from 2013/14 landslide deposit and Kick-'em-Jack hornblendes (Fig. 6a, b & c). These components are consistently lower in concentration in Kick-'em-Jack and 2013/14 landslide deposits compared to the rest of Kick-'em-Jenny samples.

### 3.2.4. Plagioclase

Phenocryst plagioclase in both Kick-'em-Jenny and Kick-'em-Jack volcanoes is predominantly anorthitic ( $\text{An}_{96-64}$ ), with only 2013/14 landslide deposits exhibiting An contents <64 (Fig. 7). Microlite compositions range  $\text{An}_{79-46}$ . Iron concentrations increase with An content until about  $\text{An}_{65}$ , after which Fe contents decrease with increasing An content. This behaviour mimics the onset of magnetite saturation observed in whole rocks (see below). Potassium concentrations decrease with increasing An content. As with other phases, there is no discernible chemical distinction between phenocrysts and glomerocrysts. The large isolated plagioclase phenocrysts have a propensity for An contents  $\geq 90$ .

There is a significant difference in the extent of normal zoning between 1972 and 2013/14 phenocrysts: 1972 plagioclase display cores ranging from  $\text{An}_{94-84}$  while plagioclase cores from 2013/14 range



**Fig. 4.** Backscattered electron images highlighting olivine textures and zoning in Kick-'em-Jenny lavas. (a) Isolated and clotted olivine with thinly zoned rims (KEJ010, 1972). (b) Isolated phenocrysts of zoned olivine, clinopyroxene, hornblende and plagioclase (KEJ018). (c) Zoned olivine clotted with overgrowth minerals of hornblende and clinopyroxene (KEJ032). (d) Zoned olivine (altered rim) and plagioclase (sieve textured) (KEJ014).

from An<sub>96–69</sub>, with rim variations of An<sub>87–68</sub> and An<sub>91–63</sub> respectively. Plagioclase phenocrysts from 2013/14 landslide deposits exhibit even wider variations in An content: cores An<sub>93–49</sub>, rims An<sub>85–38</sub>. There is much more compositional variability in 2013/14 edifice samples compared to 1972. Normally zoned Kick-'em-Jack plagioclases have cores An<sub>94–75</sub> and rims An<sub>88–65</sub> that are similar in terms of compositional variation to Kick-'em-Jenny 1972 samples. Reverse zoning was not observed in Kick-'em-Jenny 1972 or 2013/14 landslide deposits. However, in 2013/14 edifice deposits, rimward increases of up to a 9% An are observed.

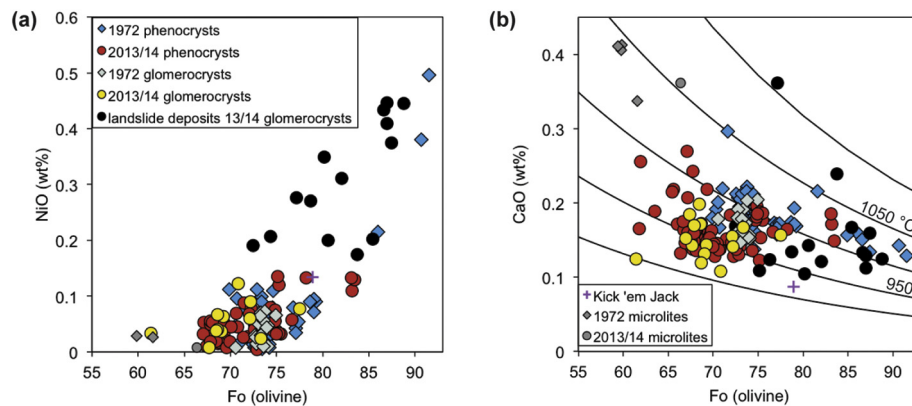
### 3.2.5. Spinel

Spinel compositions are predominantly titaniferous magnetite (Supplementary Fig. S2). 2013/14 landslide deposit and Kick-'em-Jack

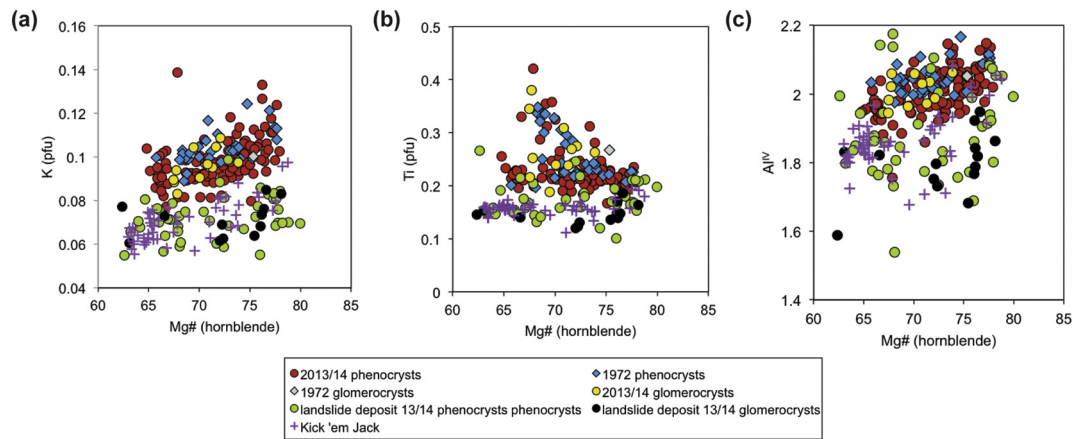
spinel are distinct from 1972 and 2013/14 edifice deposit spinels in terms of Ti content. Chrome-rich spinels are observed in a few 2013/14 landslide deposits (Cr# >60) as inclusions in hornblende, clinopyroxene and olivine (Fo<sub>82</sub>), in glomerocrysts or as isolated grains (Supplementary Fig. S2).

### 3.3. Whole-rock chemistry

Selected major and trace element variations in Kick-'em-Jenny and Kick-'em-Jack lavas are shown on MgO plots in Figs. 8 and 9. Full datasets for Kick-'em-Jenny and Kick-'em-Jack lavas are presented in Supplementary Table S3. Whole-rock data for 1972 lavas shown on Figs. 8 and 9 are taken from Sigurdsson and Shepherd (1974) and Devine and Sigurdsson (1995).



**Fig. 5.** Composition of Kick-'em-Jenny and Kick-'em-Jack olivine phenocrysts and glomerocrysts. (a) Forsterite content decreases steadily with NiO. The highest NiO contents are recorded in 1972 and 2013/14 landslide deposits. (b) Variation in CaO with Fo. Isolines calculated using the Ca-in-olivine thermometer of Shejwalkar and Coogan (2013) for orthopyroxene bearing lavas (KEJ031 and KEJ032).



**Fig. 6.** Variation in hornblende composition for Kick-'em-Jenny and Kick-'em-Jack lavas. 2013/14 landslide deposit and Kick-'em-Jack samples are broadly distinct from 1972 and 2013/14 edifice deposits.

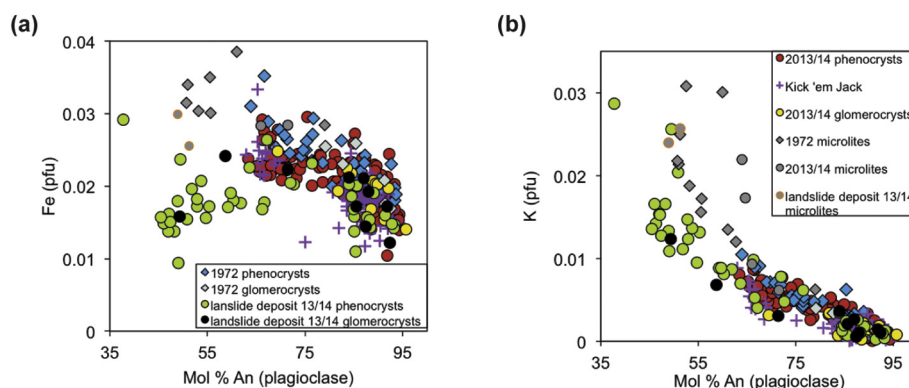
Kick-'em-Jenny lavas range from medium-K basalts to andesites (46.5–61.5 wt% SiO<sub>2</sub>) with MgO contents of 2.5 to 12.7 wt%. In general, lavas define a plausible liquid line of descent with increasing SiO<sub>2</sub>, alkalis and FeO<sub>Total</sub> with decreasing MgO and inflections in Al<sub>2</sub>O<sub>3</sub> and CaO at around 4.5 wt% MgO, consistent with the appearance of plagioclase on the liquidus (Fig. 8). However, there are distinct differences between the three datasets presented (1972, 2013/2014 and 2013/2014 landslide deposits). Lavas collected from 1972 cover a wider range of compositions from high-MgO (12.7 wt%) basalts to basaltic andesites whereas post-1972 deposits are predominantly basaltic andesite, with minor basalts. 2013/2014 lavas extend to higher Al<sub>2</sub>O<sub>3</sub> at a given MgO (Fig. 8c), than the other two types of deposits. CaO, and to a lesser extent Al<sub>2</sub>O<sub>3</sub>, show two prominent kinks at around 8 and 10 wt% MgO (Fig. 8b). Samples from 2013/14 landslide deposits produce distinct trends for SiO<sub>2</sub>, CaO, Al<sub>2</sub>O<sub>3</sub> and FeO<sub>Total</sub>. They show higher SiO<sub>2</sub> and lower CaO and Al<sub>2</sub>O<sub>3</sub> at given MgO compared to crater deposits from 1972 and 2013/14. The FeO<sub>Total</sub> is considerably lower in samples from 2013/14 landslide deposits along a trend parallel to lavas.

Lava trace elements also display fractionation trends with MgO wt%. Rb, Zr, Ba, Sr concentrations decrease with increasing MgO (Fig. 9), with 1972 and 2013/14 samples having the lowest and highest concentrations of these elements respectively. The Ba signature displays two parallel trends for 1972 and 2013/2014 crater deposits, while landslide deposits bridge these two parallel trends. Nickel and Cr (Fig. 9d and Cr in Supplementary Table S3b) concentrations decrease with decreasing MgO, with Ni reaching 232 ppm in most primitive samples. Cr contents were not measured in 1972 samples by Devine and Sigurdsson (1995). Yttrium shows an apparent maximum at around 8 wt% MgO. At lower MgO contents, Y decreases slightly, indicative of mildly compatible

behaviour during differentiation. Although samples are phenocryst-rich, mass modes of the majority of phenocryst phases do not show any obvious correlation with whole rock incompatible trace element (Zr, Rb, Ba) compositions (Supplementary Fig. S3).

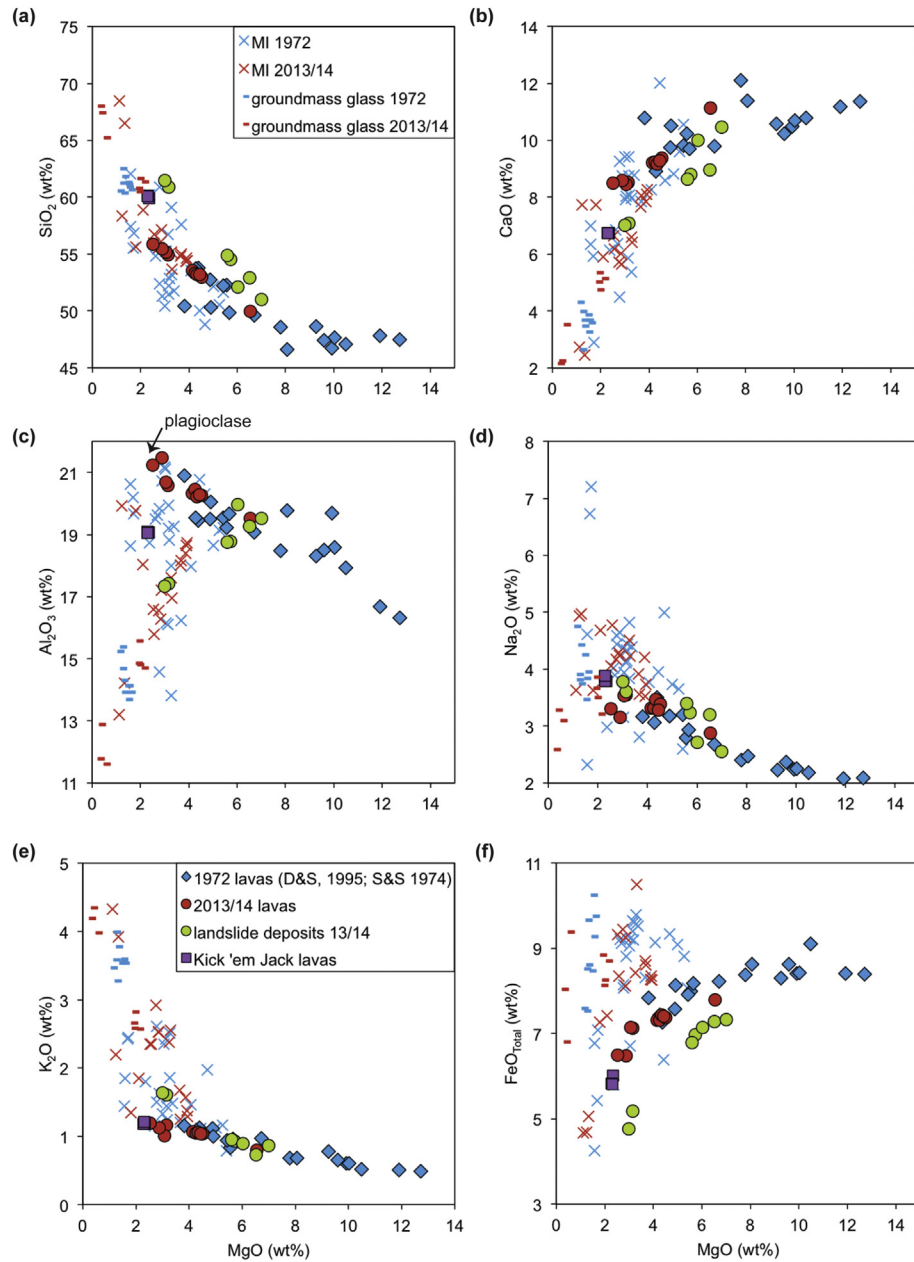
#### 3.4. Glass chemistry

Melt inclusion (MI) hosts include olivine, clinopyroxene and hornblende. Inclusions range in size from <10 to 50 μm. Only glassy, transparent MIs without daughter minerals were selected for analysis. MIs have pale brown colour in transmitted light and vary in shape from rounded to oval. While vapour bubbles are common, some MIs are bubble-free. Bubbles may contain substantial concentrations of CO<sub>2</sub>, which may lead to underestimates of the original concentration of dissolved CO<sub>2</sub> if analysing only the glass (Moore et al., 2015). MI major element compositions are reported in Table 3 and trace element chemistry in Supplementary Table S4. Post entrapment crystallisation (PEC) for MIs was up to 22% with approximately 75% of analyses having PEC corrections <10%. These corrections were based on comparisons of measured  $K_{\text{Fe-Mg}}$  values to equilibrium  $K_{\text{Fe-Mg}}$  values of 0.28 for clinopyroxene and 0.31 for hornblende hosts (from Run 3-1, Melekhova et al., 2017), with FeO = FeO<sub>Total</sub>. For olivine hosts, we used an equilibrium value of 0.22, which represents the average  $K_{\text{FeTotal-Mg}}$  calculated for Kick-'em-Jenny lavas of Sigurdsson and Shepherd (1974) for which Fe<sub>2</sub>O<sub>3</sub> and FeO were analysed separately. However, since PEC is generally small, we have only incorporated PEC corrections for olivine hosted MIs to the dataset (Table 3). We did not find any well preserved MIs from Kick-'em-Jack or Kick-'em-Jenny landslide deposits samples.



**Fig. 7.** Variation in plagioclase composition for Kick-'em-Jenny and Kick-'em-Jack lavas.





**Fig. 8.** Bulk-rock major element chemistry of Kick-'em-Jenny and Kick-'em-Jack lavas, melt inclusions and groundmass from this study, Sigurdsson and Shepherd (1974) (S&S, 1974) and Devine and Sigurdsson (1995) (D&S, 1995). Olivine-hosted melt inclusions are corrected for post-entrapment crystallisation. Vector added showing onset of plagioclase crystallisation.

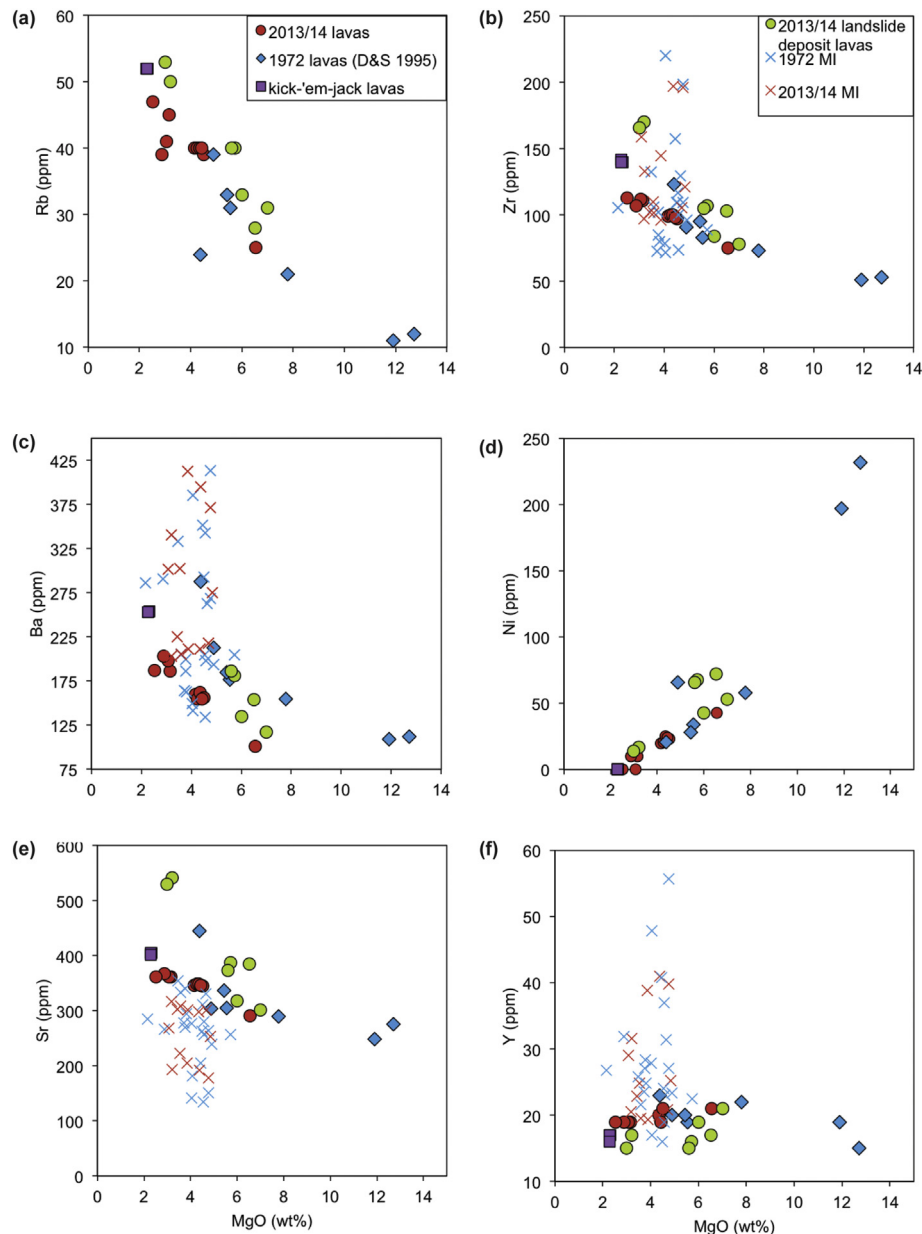
We selected twenty-seven glassy MIs ( $\geq 30 \mu\text{m}$  in diameter) from Kick-'em-Jenny 1972 lavas and seventeen from 2013/14 lavas for analysis (Table 3). MI compositions (normalized to 100% volatile-free) vary from basalt to dacite with 1.6 to 5.4 wt% MgO for 1972 and 0.5 to 3.9 wt% MgO for 2013/14 (Fig. 8). PEC-corrected MgO contents range 2.1 to 5.5 wt% for 1972 and 1.1 to 4.8 wt% for 2013/14. Chlorine contents range from 2200 to 8800 ppm; sulphur contents up to 2100 ppm. There is no compositional difference between MIs trapped in any host phase. The most evolved MIs for 1972 and 2013/14 are found in clinopyroxene and olivine hosts respectively, with those from 2013/14 having the most evolved MIs for the entire series. Some samples contain MIs with varied melt compositions, for example, basalt, basaltic andesite and andesite melts were analysed in sample KEJ010-1972 (Table 3).

Groundmass glasses (normalized to anhydrous) range from 60 to 68 wt%  $\text{SiO}_2$  with  $< 2.1$  wt% MgO (Fig. 8, Supplementary Table S5). As for MIs, 2013/14 groundmass glasses are more evolved than 1972. For

the most part, groundmass glasses and MIs follow a linear trend towards bulk rock compositions.

Volatile concentrations (measured by SIMS) range from 0.2 to 5.0 wt%  $\text{H}_2\text{O}$  and  $\leq 3157$  ppm  $\text{CO}_2$ . Volatile contents are only minimally affected by PEC corrections. The highest  $\text{CO}_2$  contents are found in hornblende and olivine hosts (Fig. 12). MIs contained in clinopyroxene and hornblende tend to have higher  $\text{H}_2\text{O}$  concentrations ( $> 4$  wt%) than those in olivine. Olivine hosts contain  $\text{Fo}_{87-65}$ .  $\text{H}_2\text{O}$  correlates fairly well with some major elements: FeO,  $\text{K}_2\text{O}$ ,  $\text{Na}_2\text{O}$  and  $\text{TiO}_2$  generally decrease with increasing  $\text{H}_2\text{O}$ , while CaO and  $\text{Al}_2\text{O}_3$  increase. The variation in  $\text{H}_2\text{O}$  with  $\text{SiO}_2$  shows trends consistent with decompression crystallisation and syn-eruptive degassing as described by Blundy and Cashman (2005) (Fig. 10c).

Most trace elements (B, Ti, Y, Zr, Nb, Ba, Ce) show steady low concentrations at high  $\text{H}_2\text{O}$  contents (Supplementary Table S4 and Fig. 10a) with sudden inflections below  $\sim 5$  wt% MgO (Fig. 9a, b, c & f), similar to the behaviour of  $\text{K}_2\text{O}$  (Fig. 8e). Strontium shows the opposite behaviour



**Fig. 9.** Bulk-rock trace element chemistry of Kick-'em-Jenny and Kick-'em-Jack lavas and melt inclusions from this study and Devine and Sigurdsson (1995) (D&S, 1995).

(Fig. 10b), with a marked drop in concentration below 5 wt% MgO (Fig. 9e), similar to  $\text{Al}_2\text{O}_3$  (Fig. 8a), and consistent with the onset of plagioclase crystallisation.

### 3.5. Intensive parameters

#### 3.5.1. Thermometry of lavas

Mineral, MI and groundmass glass compositions were used to determine pre-eruptive magma storage conditions (P-T-H<sub>2</sub>O-fO<sub>2</sub>) for Kick-'em-Jenny and Kick-'em-Jack. A combination of geothermometers was used to estimate equilibrium temperatures of phenocryst assemblages: hornblende-plagioclase (Holland and Blundy, 1994), clinopyroxene-orthopyroxene (Brey and Kohler, 1990; Putirka, 2008) and clinopyroxene-liquid (Putirka, 2008). We chose three MIs with *Kd* values indicating that PEC effects were minimal for the clinopyroxene-liquid calculations. Major element data for touching rims and/or included phases were used in hornblende-plagioclase calculations to ensure a close approach to textural equilibrium in zoned phases. Because pressure has been shown to have a modest

dependence on thermometry ( $\pm 72^\circ\text{C}/\text{GPa}$ ) at sub-volcanic pressures (Blundy and Cashman, 2008), a nominal value of 500 MPa was applied. The hornblende-plagioclase thermometer requires plagioclase to be less calcic than An<sub>90</sub> (Holland and Blundy, 1994). Both Kick-'em-Jenny and Kick-'em-Jack contained plagioclase with higher anorthite contents, however only temperatures calculated for An < 90% are included in Table 4. Clinopyroxene and orthopyroxene grains were never adjacent, therefore all combinations of individual clinopyroxene and orthopyroxene rims were used for calculations and the resulting temperatures averaged. Standard deviations for these calculations were <13 °C for KEJ031, <17 °C for KEJ032 and <59 °C for KEJ008a. Estimates of magmatic temperatures at lower pressures (i.e. eruption temperatures) were obtained using groundmass olivine-liquid assemblages (Putirka, 2008). Again, temperatures were averaged from all combinations of olivine and glass in each sample (standard deviations: 10 °C for KEJ018, 22 °C for KEJ005, 17 °C for KEJ010 (1972) and 7 °C for KEJ019).

For Kick-'em-Jenny, the temperature range obtained using the hornblende-plagioclase thermometer is lower for 1972 crater lavas



**Table 3**  
Kick-'em-Jenny phenocryst-hosted melt inclusion compositions.

| Sample                         | KEJ010       | KEJ010        | KEJ010        | KEJ010        | KEJ010       | KEJ010       | KEJ005    | KEJ005     | KEJ005      | KEJ005     | KEJ018     |
|--------------------------------|--------------|---------------|---------------|---------------|--------------|--------------|-----------|------------|-------------|------------|------------|
| Name                           | kej010mi2ats | kej010mi10ats | kej010mi10bts | kej010mi10cts | kej010mi16ts | kej010mi22ts | kej005mi5 | kej005mi20 | kej005mi21b | kej005mi9  | kej018mi1b |
| Host                           | ol           | ol            | ol            | ol            | ol           | ol           | ol        | ol         | ol          | ol         | ol         |
| Period                         | 1972         | 1972          | 1972          | 1972          | 1972         | 1972         | 1972      | 1972       | 1972        | 1972       | 1972       |
| SiO <sub>2</sub>               | 60.46        | 50.86         | 56.71         | 55.44         | 59.99        | 52.19        | 51.85     | 52.11      | 51.40       | 50.62      | 49.30      |
| TiO <sub>2</sub>               | 1.87         | 1.21          | 1.74          | 1.84          | 2.18         | 1.24         | 1.41      | 1.13       | 1.36        | 1.40       | 1.20       |
| Al <sub>2</sub> O <sub>3</sub> | 14.73        | 19.19         | 16.54         | 16.64         | 14.42        | 19.19        | 19.97     | 17.88      | 19.44       | 19.99      | 21.11      |
| FeO                            | 7.50         | 9.05          | 8.89          | 8.53          | 9.10         | 8.70         | 7.63      | 9.05       | 8.60        | 8.63       | 5.95       |
| MnO                            | 0.14         | 0.23          | 0.27          | 0.21          | 0.23         | 0.20         | 0.19      | 0.19       | 0.15        | 0.11       | 0.16       |
| MgO                            | 1.80         | 2.67          | 1.74          | 1.68          | 1.73         | 1.58         | 1.91      | 2.55       | 1.79        | 2.23       | 2.80       |
| CaO                            | 4.53         | 8.75          | 6.02          | 8.28          | 5.59         | 8.23         | 9.32      | 7.91       | 8.85        | 9.44       | 12.20      |
| Na <sub>2</sub> O              | 4.64         | 4.38          | 4.23          | 4.06          | 4.01         | 4.52         | 4.46      | 4.78       | 4.39        | 4.17       | 4.01       |
| K <sub>2</sub> O               | 2.64         | 1.47          | 2.49          | 2.42          | 2.62         | 1.59         | 1.51      | 1.85       | 1.46        | 1.24       | 1.13       |
| P <sub>2</sub> O <sub>5</sub>  | 0.41         | 0.18          | 0.18          | 0.21          | 0.34         | 0.11         | 0.16      | 0.14       | 0.15        | 0.08       | 0.18       |
| SO <sub>2</sub>                | bdl          | 0.07          | bdl           | 0.07          | 0.05         | 0.07         | 0.07      | 0.04       | 0.11        | 0.04       | 0.08       |
| Cl                             | 0.60         | 0.40          | 0.66          | 0.33          | 0.68         | 0.40         | 0.46      | 0.45       | 0.41        | 0.38       | 0.39       |
| Cr <sub>2</sub> O <sub>3</sub> | bdl          | bdl           | bdl           | bdl           | bdl          | bdl          | bdl       | bdl        | bdl         | bdl        | bdl        |
| H <sub>2</sub> O               | 0.53         | 0.85          | 0.45          | 0.65          | 0.50         | 1.68         | 0.38      | 1.04       | 1.31        | 1.36       | 0.72       |
| sd                             | 0.01         | 0.02          | 0.01          | 0.01          | 0.01         | 0.05         | 0.01      | 0.02       | 0.04        | 0.03       | 0.02       |
| CO <sub>2</sub>                | 93           | bdl           | bdl           | 43            | bdl          | 221          | 67        | 204        | 137         | 834        | 459        |
| sd                             | 9            | –             | –             | 10            | –            | 16           | 17        | 50         | 24          | 120        | 63         |
| P <sub>sat</sub>               | 19           | 10            | 3             | 13            | 4            | 72           | 16        | 49         | 47          | 205        | 125        |
| Total                          | 99.85        | 99.31         | 99.92         | 100.36        | 101.44       | 99.70        | 99.35     | 99.12      | 99.43       | 99.69      | 99.23      |
| Restored compositions          |              |               |               |               |              |              |           |            |             |            |            |
| SiO <sub>2</sub>               | 59.87        | 50.63         | 55.97         | 54.77         | 59.03        | 51.57        | 51.52     | 51.84      | 50.91       | 50.29      | 48.94      |
| TiO <sub>2</sub>               | 1.82         | 1.19          | 1.68          | 1.77          | 2.08         | 1.18         | 1.38      | 1.11       | 1.31        | 1.37       | 1.16       |
| Al <sub>2</sub> O <sub>3</sub> | 14.34        | 18.85         | 15.89         | 16.00         | 13.79        | 18.36        | 19.50     | 17.55      | 18.73       | 19.50      | 20.33      |
| FeO                            | 7.93         | 9.31          | 9.51          | 9.14          | 9.77         | 9.38         | 8.01      | 9.33       | 9.17        | 9.00       | 6.25       |
| MnO                            | 0.15         | 0.24          | 0.28          | 0.22          | 0.24         | 0.21         | 0.20      | 0.19       | 0.16        | 0.11       | 0.16       |
| MgO                            | 2.73         | 3.31          | 3.09          | 3.03          | 3.28         | 3.12         | 2.73      | 3.20       | 3.08        | 3.10       | 4.34       |
| CaO                            | 4.42         | 8.60          | 5.79          | 7.97          | 5.36         | 7.88         | 9.11      | 7.76       | 8.53        | 9.21       | 11.75      |
| Na <sub>2</sub> O              | 4.52         | 4.30          | 4.06          | 3.91          | 3.84         | 4.33         | 4.36      | 4.69       | 4.23        | 4.06       | 3.86       |
| K <sub>2</sub> O               | 2.57         | 1.45          | 2.39          | 2.33          | 2.51         | 1.52         | 1.48      | 1.81       | 1.41        | 1.21       | 1.09       |
| Cr <sub>2</sub> O <sub>3</sub> | 0.00         | 0.00          | 0.00          | 0.00          | 0.00         | 0.00         | 0.02      | 0.00       | 0.01        | 0.00       | 0.00       |
| Sample                         | KEJ005       | KEJ010        | KEJ010        | KEJ010        | KEJ010       | KEJ010       | KEJ018    | KEJ009     | KEJ009      | KEJ010     | KEJ009     |
| Name                           | kej005mi21a  | kej010mi6ts   | kej010mi12ts  | kej010mi7     | kej010mi9    | kej010mi8    | kej018mi3 | kej009mi3a | kej009mi3b  | kej010mi6a | kej009mi6  |
| Host                           | ol           | ol            | cpx           | cpx           | cpx          | cpx          | cpx       | cpx        | cpx         | cpx        | hbl        |
| Period                         | 1972         | 1972          | 1972          | 1972          | 1972         | 1972         | 1972      | 1972       | 1972        | 1972       | 1972       |
| SiO <sub>2</sub>               | 50.63        | 49.61         | 56.53         | 58.07         | 48.08        | 56.05        | 53.02     | 54.20      | 54.06       | 53.96      | 49.80      |
| TiO <sub>2</sub>               | 1.13         | 1.10          | 1.21          | 0.51          | 0.96         | 1.02         | 1.01      | 0.80       | 0.89        | 1.18       | 1.02       |
| Al <sub>2</sub> O <sub>3</sub> | 20.94        | 21.23         | 19.57         | 19.30         | 17.92        | 17.44        | 17.19     | 19.26      | 19.13       | 15.22      | 18.86      |
| FeO                            | 8.12         | 8.47          | 7.04          | 3.98          | 7.52         | 5.39         | 6.26      | 9.11       | 8.89        | 7.78       | 8.68       |
| MnO                            | 0.11         | 0.14          | 0.16          | 0.15          | 0.19         | 0.12         | 0.19      | 0.18       | 0.20        | 0.28       | 0.25       |
| MgO                            | 1.76         | 2.09          | 1.71          | 1.47          | 5.05         | 2.20         | 1.46      | 2.59       | 2.63        | 3.45       | 5.18       |
| CaO                            | 8.84         | 9.44          | 2.89          | 5.94          | 9.82         | 5.71         | 6.46      | 6.77       | 6.20        | 7.36       | 9.47       |
| Na <sub>2</sub> O              | 4.69         | 4.38          | 7.16          | 2.17          | 2.43         | 2.78         | 4.25      | 3.97       | 3.79        | 2.63       | 3.60       |
| K <sub>2</sub> O               | 1.64         | 1.33          | 2.43          | 1.35          | 0.73         | 1.67         | 1.71      | 1.16       | 1.12        | 1.12       | 1.15       |
| P <sub>2</sub> O <sub>5</sub>  | 0.20         | 0.14          | 0.20          | 0.09          | 0.14         | 0.28         | 0.22      | 0.22       | 0.20        | 0.20       | 0.14       |
| SO <sub>2</sub>                | 0.12         | 0.11          | 0.06          | 0.19          | 0.13         | 0.05         | 0.05      | 0.08       | 0.07        | 0.14       | 0.05       |
| Cl                             | 0.37         | 0.37          | 0.48          | 0.34          | 0.22         | 0.42         | 0.44      | 0.43       | 0.35        | 0.33       | 0.32       |
| Cr <sub>2</sub> O <sub>3</sub> | bdl          | bdl           | bdl           | bdl           | bdl          | bdl          | bdl       | bdl        | bdl         | bdl        | bdl        |

(continued on next page)

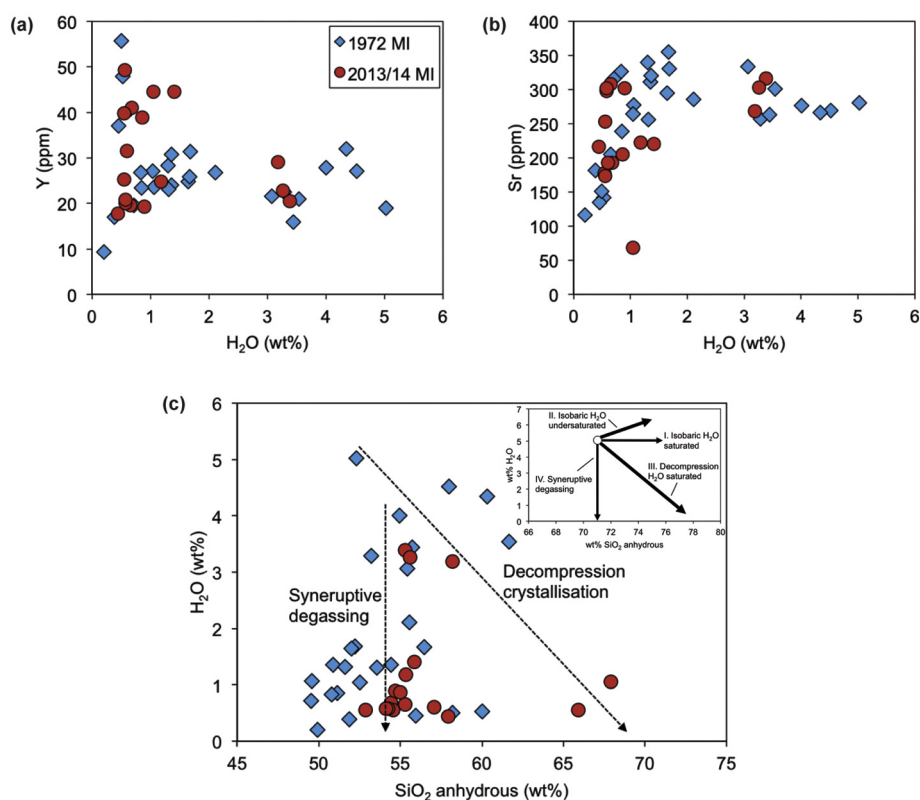
Table 3 (continued)

| Sample                         | KEJ005      | KEJ010      | KEJ010       | KEJ010      | KEJ010    | KEJ010     | KEJ018     | KEJ009     | KEJ009     | KEJ010      | KEJ009      |
|--------------------------------|-------------|-------------|--------------|-------------|-----------|------------|------------|------------|------------|-------------|-------------|
| Name                           | kej005mi21a | kej010mi6ts | kej010mi12ts | kej010mi7   | kej010mi9 | kej010mi8  | kej018mi3  | kej009mi3a | kej009mi3b | kej010mi6a  | kej009mi6   |
| Host                           | ol          | ol          | cpx          | cpx         | cpx       | cpx        | cpx        | cpx        | cpx        | cpx         | hbl         |
| Period                         | 1972        | 1972        | 1972         | 1972        | 1972      | 1972       | 1972       | 1972       | 1972       | 1972        | 1972        |
| H <sub>2</sub> O               | 0.83        | 0.20        | 1.67         | 3.54        | 5.02      | 4.34       | 3.44       | 1.35       | 3.07       | 4.53        | 1.65        |
| sd                             | 0.02        | 0.02        | 0.04         | 0.20        | 0.11      | 0.12       | 0.09       | 0.03       | 0.12       | 0.10        | 0.03        |
| CO <sub>2</sub>                | 666         | 56          | bdl          | 439         | 784       | 694        | 75         | bdl        | bdl        | 767         | bdl         |
| sd                             | 33          | 21          | –            | 43          | 29        | 57         | 40         | –          | –          | 38          | –           |
| P <sub>sat</sub>               | 164         | 15          | 31           | 185         | 325       | 258        | 105        | 21         | 87         | 273         | 35          |
| Total                          | 99.38       | 98.61       | 101.11       | 97.10       | 98.21     | 97.47      | 95.74      | 100.17     | 100.40     | 98.21       | 100.19      |
| Restored compositions          |             |             |              |             |           |            |            |            |            |             |             |
| SiO <sub>2</sub>               | 50.23       | 49.34       | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| TiO <sub>2</sub>               | 1.09        | 1.07        | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| Al <sub>2</sub> O <sub>3</sub> | 20.30       | 20.71       | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| FeO                            | 8.63        | 8.86        | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| MnO                            | 0.12        | 0.15        | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| MgO                            | 2.82        | 2.95        | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| CaO                            | 8.58        | 9.21        | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| Na <sub>2</sub> O              | 4.55        | 4.27        | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| K <sub>2</sub> O               | 1.59        | 1.30        | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| Cr <sub>2</sub> O <sub>3</sub> | 0.00        | 0.00        | –            | –           | –         | –          | –          | –          | –          | –           | –           |
| Sample                         | KEJ009      | KEJ011      | KEJ011       | KEJ005      | KEJ009    | KEJ050     | KEJ042     | KEJ042     | KEJ019     | KEJ019      | KEJ019      |
| Name                           | kej009mi4   | kej011mi3a  | kej011mi3b   | kej005mi22b | kej009mi5 | kej050mi16 | kej042mi13 | kej042mi9  | kej019mi20 | kej019mi22a | kej019mi21b |
| Host                           | hbl         | hbl         | hbl          | hbl         | hbl       | ol         | ol         | ol         | ol         | hbl         | hbl         |
| Period                         | 1972        | 1972        | 1972         | 1972        | 1972      | 2013       | 2013       | 2013       | 2014       | 2014        | 2014        |
| SiO <sub>2</sub>               | 48.05       | 51.62       | 51.52        | 54.30       | 52.15     | 65.84      | 67.11      | 57.89      | 57.21      | 54.39       | 55.49       |
| TiO <sub>2</sub>               | 0.82        | 0.98        | 0.93         | 0.98        | 0.94      | 1.47       | 1.73       | 0.84       | 0.89       | 0.62        | 2.19        |
| Al <sub>2</sub> O <sub>3</sub> | 20.00       | 17.32       | 19.67        | 19.76       | 18.64     | 14.14      | 13.07      | 20.03      | 17.55      | 19.34       | 15.44       |
| FeO                            | 9.19        | 8.80        | 6.26         | 5.31        | 9.09      | 4.75       | 4.09       | 4.16       | 7.19       | 7.11        | 8.17        |
| MnO                            | 0.12        | 0.26        | 0.12         | 0.18        | 0.21      | 0.09       | 0.09       | 0.08       | 0.19       | 0.13        | 0.21        |
| MgO                            | 4.61        | 3.93        | 2.82         | 1.64        | 5.00      | 0.10       | 0.53       | 0.58       | 1.99       | 1.76        | 2.52        |
| CaO                            | 8.45        | 7.98        | 7.39         | 5.79        | 8.82      | 2.44       | 2.70       | 7.78       | 5.75       | 7.57        | 6.05        |
| Na <sub>2</sub> O              | 4.91        | 3.63        | 2.93         | 6.58        | 3.74      | 4.95       | 3.59       | 4.97       | 4.55       | 3.55        | 4.67        |
| K <sub>2</sub> O               | 1.94        | 1.41        | 0.99         | 2.38        | 1.12      | 3.91       | 4.28       | 2.20       | 1.80       | 1.32        | 2.30        |
| P <sub>2</sub> O <sub>5</sub>  | 0.10        | 0.14        | 0.10         | 0.23        | 0.11      | 0.51       | 0.53       | 0.28       | 0.13       | 0.21        | 0.23        |
| SO <sub>2</sub>                | bdl         | 0.05        | 0.19         | 0.06        | bdl       | bdl        | 0.04       | 0.06       | 0.04       | 0.03        | 0.03        |
| Cl                             | 0.24        | 0.32        | 0.24         | 0.59        | 0.22      | 0.76       | 0.87       | 0.57       | 0.51       | 0.32        | 0.57        |
| Cr <sub>2</sub> O <sub>3</sub> | bdl         | bdl         | bdl          | bdl         | bdl       | bdl        | bdl        | bdl        | bdl        | bdl         | bdl         |
| H <sub>2</sub> O               | 1.06        | 3.29        | 4.01         | 2.11        | 1.30      | 0.56       | 1.05       | 0.44       | 3.19       | 3.39        | 1.41        |
| sd                             | 0.02        | 0.10        | 0.25         | 0.05        | 0.04      | 0.02       | 0.03       | 0.01       | 0.07       | 0.09        | 0.03        |
| CO <sub>2</sub>                | bdl         | 246         | 663          | 145         | bdl       | 2513       | 854        | 311        | 165        | 102         | 3157        |
| sd                             | –           | 20          | 76           | 7           | –         | 250        | 154        | 81         | 52         | 71          | 847         |
| P <sub>sat</sub>               | 15          | 133         | 253          | 66          | 20        | 552        | 164        | 91         | 114        | 116         | 513         |
| Total                          | 99.54       | 99.73       | 97.18        | 99.91       | 101.35    | 99.52      | 99.68      | 99.88      | 101.01     | 99.74       | 99.28       |
| Restored compositions          |             |             |              |             |           |            |            |            |            |             |             |
| SiO <sub>2</sub>               | –           | –           | –            | –           | –         | 65.56      | 66.57      | 57.50      | 57.19      | –           | –           |
| TiO <sub>2</sub>               | –           | –           | –            | –           | –         | 1.45       | 1.70       | 0.83       | 0.89       | –           | –           |
| Al <sub>2</sub> O <sub>3</sub> | –           | –           | –            | –           | –         | 14.00      | 12.83      | 19.64      | 17.53      | –           | –           |
| FeO                            | –           | –           | –            | –           | –         | 4.98       | 4.54       | 4.63       | 7.21       | –           | –           |
| MnO                            | –           | –           | –            | –           | –         | 0.10       | 0.10       | 0.09       | 0.19       | –           | –           |
| MgO                            | –           | –           | –            | –           | –         | 1.32       | 1.09       | 1.23       | 2.04       | –           | –           |
| CaO                            | –           | –           | –            | –           | –         | 2.42       | 2.66       | 7.63       | 5.74       | –           | –           |
| Na <sub>2</sub> O              | –           | –           | –            | –           | –         | 4.90       | 3.53       | 4.87       | 4.54       | –           | –           |

|                                |             |             |             |             |             |             |             |             |            |             |             |
|--------------------------------|-------------|-------------|-------------|-------------|-------------|-------------|-------------|-------------|------------|-------------|-------------|
| K <sub>2</sub> O               | –           | –           | –           | –           | –           | 3.87        | 4.21        | 2.16        | 1.79       | –           | –           |
| Cr <sub>2</sub> O <sub>3</sub> | –           | –           | –           | –           | –           | 0.00        | 0.00        | 0.00        | 0.00       | –           | –           |
| Sample                         | KEJ019      | KEJ019      | KEJ019      | KEJ019      | KEJ019      | KEJ019      | KEK019      | KEJ019      | KEJ019     | KEJ019      | KEJ019      |
| Name                           | kej019mi22b | kej019mi22c | kej019mi23a | kej019mi23b | kej019mi23d | kej019mi21a | kej019mi21c | kej019mi21d | kej019mi19 | kej019mi23c | kej019mi22d |
| Host                           | hbl         | hbl         | hbl         | hbl         | hbl         | hbl         | hbl         | hbl         | hbl        | hbl         | hbl         |
| Period                         | 2014        | 2014        | 2014        | 2014        | 2014        | 2014        | 2014        | 2014        | 2014       | 2014        | 2014        |
| SiO <sub>2</sub>               | 55.16       | 55.71       | 54.16       | 53.71       | 53.48       | 56.78       | 53.95       | 53.10       | 53.29      | 54.45       | 53.42       |
| TiO <sub>2</sub>               | 1.63        | 1.62        | 0.71        | 0.73        | 0.72        | 1.40        | 1.01        | 0.83        | 1.29       | 0.77        | 1.32        |
| Al <sub>2</sub> O <sub>3</sub> | 16.43       | 16.17       | 18.49       | 18.50       | 18.10       | 17.09       | 17.23       | 17.53       | 16.82      | 17.89       | 15.85       |
| FeO                            | 9.38        | 9.19        | 8.27        | 8.15        | 8.13        | 8.06        | 8.23        | 8.31        | 10.42      | 8.64        | 8.90        |
| MnO                            | 0.21        | 0.22        | 0.21        | 0.19        | 0.19        | 0.19        | 0.20        | 0.27        | 0.22       | 0.24        | 0.19        |
| MgO                            | 2.72        | 2.82        | 3.88        | 3.88        | 3.81        | 2.84        | 3.19        | 3.56        | 3.29       | 3.63        | 2.41        |
| CaO                            | 5.68        | 5.98        | 8.03        | 8.16        | 7.72        | 5.62        | 6.46        | 7.78        | 6.36       | 7.60        | 6.35        |
| Na <sub>2</sub> O              | 4.14        | 4.25        | 3.49        | 3.69        | 4.14        | 4.21        | 4.40        | 3.44        | 4.20       | 3.89        | 3.87        |
| K <sub>2</sub> O               | 2.90        | 2.51        | 1.36        | 1.27        | 1.54        | 2.51        | 2.33        | 1.20        | 2.53       | 1.66        | 2.25        |
| P <sub>2</sub> O <sub>5</sub>  | 0.41        | 0.32        | 0.18        | 0.15        | 0.15        | 0.30        | 0.20        | 0.13        | 0.19       | 0.15        | 0.27        |
| SO <sub>2</sub>                | bdl         | 0.03        | bdl         | bdl         | 0.03        | bdl         | 0.03        | bdl         | bdl        | bdl         | bdl         |
| Cl                             | 0.62        | 0.52        | 0.33        | 0.31        | 0.33        | 0.34        | 0.56        | 0.37        | 0.66       | 0.36        | 0.65        |
| Cr <sub>2</sub> O <sub>3</sub> | bdl         | bdl         | bdl         | bdl         | bdl         | bdl         | bdl         | bdl         | bdl        | bdl         | bdl         |
| H <sub>2</sub> O               | 0.68        | 0.55        | 0.66        | 0.58        | 0.90        | 0.60        | 1.18        | 3.26        | 0.55       | 0.58        | 0.87        |
| sd                             | 0.02        | 0.01        | 0.01        | 0.02        | 0.02        | 0.02        | 0.03        | 0.07        | 0.01       | 0.01        | 0.02        |
| CO <sub>2</sub>                | 109         | 48          | 98          | 104         | 94          | 90          | 91          | 52          | 37         | 67          | 250         |
| sd                             | 63          | 41          | 96          | 81          | 18          | 51          | 21          | 9           | 14         | 15          | 108         |
| P <sub>sat</sub>               | 25          | 13          | 27          | 27          | 29          | 22          | 33          | 98          | 11         | 18          | 70          |
| Total                          | 99.96       | 99.89       | 99.81       | 99.33       | 99.25       | 99.96       | 98.97       | 99.82       | 99.84      | 99.87       | 96.37       |

Major elements (in wt%) determined by electron microprobe; H<sub>2</sub>O (in wt%) and CO<sub>2</sub> (in ppm) measured by ion-microprobe; sd (standard deviation) calculated by full error propagation of counting statistics, instrument background and calibration working curve; P<sub>sat</sub> (saturation pressure) in MPa using method of [Ghiorso and Gualda \(2015\)](#); Olivine hosted melt inclusion compositions recalculated using average  $Kd_{FeTotal-Mg}$  value of 0.22 for Kick-'em-Jenny lavas of [Sigurdsson and Shepherd \(1974\)](#) for which Fe<sub>2</sub>O<sub>3</sub> and FeO were analysed separately; ol, olivine; cpx, clinopyroxene; hbl, hornblende; bdl; below detection limit.





**Fig. 10.** Variation of selected trace elements (a) Y and (b) Sr with H<sub>2</sub>O (all measured by SIMS) for Kick-'em-Jenny melt inclusions. (c) H<sub>2</sub>O vs. SiO<sub>2</sub> (normalized to 100% H<sub>2</sub>O free) plot. The dominant trends are consistent with decompression crystallisation (trend III) and syn-eruption degassing (trend IV) as displayed in the inserted schematic H<sub>2</sub>O-SiO<sub>2</sub> plot showing trends anticipated for different crystallisation mechanisms after Blundy and Cashman (2005).

(964–968 °C) than for 2013/14 crater lavas (974–1055 °C) (Table 4, Fig. 11). Clinopyroxene-liquid thermometry for 1972 samples range 976–1008 °C, which is notably wider than temperatures obtained using hornblende-plagioclase calculations. 2013/14 landslide deposit basaltic andesite temperatures range 900–975 °C using hornblende-plagioclase and 969–1012 °C using clinopyroxene-orthopyroxene thermometers. The sole andesite (KEJ008a) ranges 820–840 °C using clinopyroxene-orthopyroxene. These distinctions are clearly represented in Fig. 11. Eruption temperatures using groundmass glass and olivine microlites range 1006–1022 °C for 1972 and 1079 °C for 2014

deposits. This matches well with extrapolated temperature estimates for microlite olivine (~1025–1050 °C) using the algorithm of Shejwalkar and Coogan (2013) (Fig. 5a). Kick-'em-Jack lavas range 946–1007 °C using hornblende-plagioclase phenocryst pairs.

Mean magma temperatures are notably low (~1000 ± 30 °C) for such magnesian lavas. However, the consistency of estimates for from the range of methods gives us confidence in our temperatures. The similarity of phenocryst and eruption temperatures is consistent with rapid ascent of magmas before eruption. We will return to the implications of this observation in a later section.

**Table 4**

Temperature estimates for Kick-'em-Jenny and Kick-'em-Jack volcanoes.

| Volcano        | Sample               | T (°C)                                |                            |                            |                                    |                                        |
|----------------|----------------------|---------------------------------------|----------------------------|----------------------------|------------------------------------|----------------------------------------|
|                |                      | hbl-plag<br>Holland and Blundy (1994) | cpx-liq<br>(Putirka, 2008) | cpx-opx<br>(Putirka, 2008) | cpx-opx<br>(Brey and Kohler, 1990) | ol-liq <sup>a</sup><br>(Putirka, 2008) |
| Kick-'em-Jenny | KEJ018 (1972)        | 964                                   | –                          | –                          | –                                  | 1022                                   |
|                | KEJ005 (1972)        | 966                                   | –                          | –                          | –                                  | 1023                                   |
|                | KEJ010 (1972)        | 968                                   | 976–1008                   | –                          | –                                  | 1006                                   |
|                | KEJ009 (1972)        | –                                     | 1008                       | –                          | –                                  | –                                      |
|                | KEJ014               | 990–1010                              | –                          | –                          | –                                  | –                                      |
|                | KEJ019               | 983                                   | –                          | –                          | –                                  | 1079                                   |
|                | KEJ008               | 977–1055                              | –                          | –                          | –                                  | –                                      |
|                | KEJ048               | 974                                   | –                          | –                          | –                                  | –                                      |
|                | KEJ021               | 989                                   | –                          | –                          | –                                  | –                                      |
|                | KEJ032 <sup>b</sup>  | 920–945                               | –                          | 969–991                    | 981–1012                           | –                                      |
|                | KEJ031 <sup>b</sup>  | 900–975                               | –                          | 973–986                    | 983–1002                           | –                                      |
|                | KEJ008a <sup>b</sup> | –                                     | –                          | 840                        | 820                                | –                                      |
| Kick-'em-Jack  | KEJ082               | 987                                   | –                          | –                          | –                                  | –                                      |
|                | KEJ086               | 946–1007                              | –                          | –                          | –                                  | –                                      |
|                | KEJ084               | 955                                   | –                          | –                          | –                                  | –                                      |
|                | KEJ085               | 997                                   | –                          | –                          | –                                  | –                                      |

<sup>a</sup> Microlite-glass data.

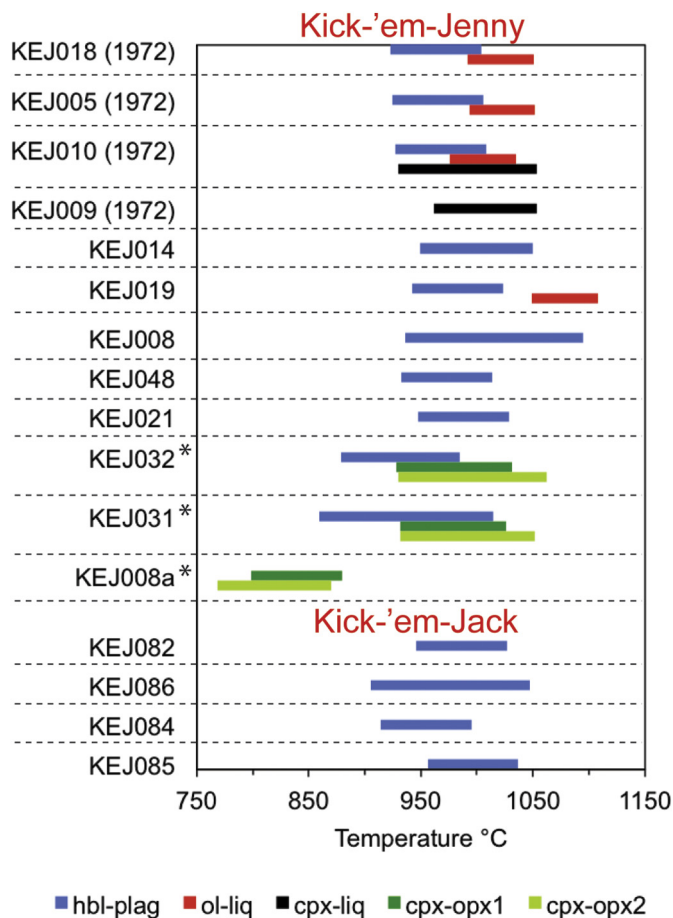
<sup>b</sup> 2013/14 landslide deposits.

### 3.5.2. Magma storage pressures

Pressures were estimated by applying the MagmaSat solubility model of [Ghiorso and Gualda \(2015\)](#) to H<sub>2</sub>O and CO<sub>2</sub> contents from phenocryst MIs assuming volatile saturation ([Fig. 12](#)). H<sub>2</sub>O and CO<sub>2</sub> concentrations in MIs and volatile saturation pressure isobars for a median MI composition are plotted in [Fig. 12a](#). A degassing trend is observed from high to low volatile contents. Magma storage pressures for Kick-'em-Jenny range up to 552 MPa, but the majority reach 325 MPa ([Fig. 12b](#), [Table 3](#)). Highest pressures were obtained from melt inclusions containing high CO<sub>2</sub> (2500–3200 ppm) with correspondingly low H<sub>2</sub>O (0.6–1.4 wt%) contents. We view these analyses with caution given the possibility that they may result from post-entrapment H<sub>2</sub>O-loss, although we cannot rule out fluxing of previously degassed magmas by CO<sub>2</sub>-rich fluids ([Caricchi et al., 2018](#)).

### 3.5.3. Oxygen fugacity

In the absence of appropriate oxybarometers to calculate the oxygen fugacity of Kick-'em-Jenny lavas, we apply the algorithm of [Kress and Carmichael \(1991\)](#) to eleven 1972 whole-rock samples from [Sigurdsson and Shepherd \(1974\)](#) who provide measured FeO/Fe<sub>2</sub>O<sub>3</sub> ratio from which we can compute the  $f_{O_2}$  corresponding to Fe-redox equilibrium at P and T. The mean  $Fe^{3+}/Fe^{Total}$  for these samples is  $0.28 \pm 0.03$ , with no clear dependence on MgO. Using P = 300 MPa and T = 950 °C,  $f_{O_2}$  estimates for eleven 1972 basalts are in the range NNO +0.5 to +1.4 log units (with a single exception at NNO +0.05).



**Fig. 11.** Temperature estimates for Kick-'em-Jenny and Kick-'em-Jack lavas. Temperature ranges include errors associated with thermometers. Landslide deposits for 2013/14 Kick-'em-Jenny samples are given the symbol \*. Thermometers used: hbl-plag, [Holland and Blundy \(1994\)](#) ( $\pm 40$  °C); ol-liq, [Putirka \(2008\)](#) ( $\pm 29$  °C); cpx-liq, [Putirka \(2008\)](#) ( $\pm 45$  °C); cpx-opx1, [Putirka \(2008\)](#) ( $\pm 40$  °C); cpx-opx2, [Brey and Kohler \(1990\)](#) ( $\pm 50$  °C).

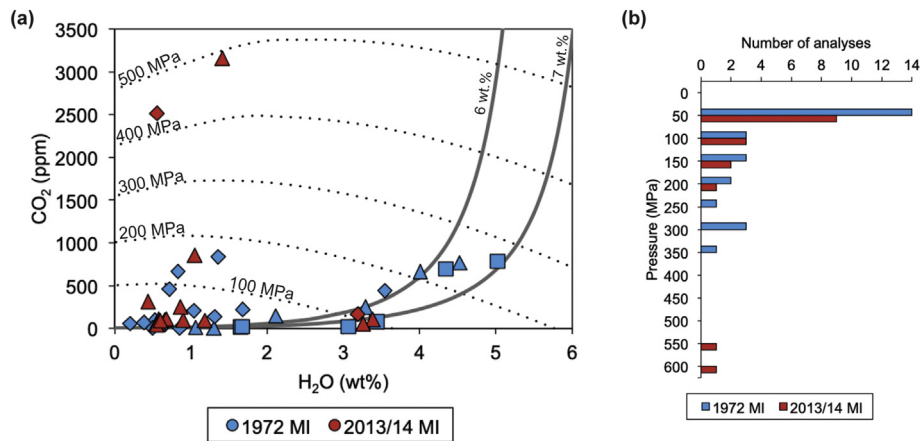
## 4. Discussion

### 4.1. Crystal cargo origins

Crystal populations in Kick-'em-Jenny lavas contain phase assemblages and textures that suggest a complex magmatic history, including both changes in intensive variables during crystallisation and incorporation of xenocryst or antecryst material. Some melt inclusions (MIs) have compositions that overlap with Kick-'em-Jenny whole rocks, although most MIs are displaced from the whole-rock trend suggesting a change in crystallisation assemblage. Kinks in major element whole rock trends for CaO and Al<sub>2</sub>O<sub>3</sub> ([Fig. 8](#)) also suggest changes in the crystallisation assemblage. Inferring the relative order of crystallisation for each phenocryst phase from included and excluded phases consequently proved problematic. Magnetite-rich (and Cr-rich when present) spinels are included in all phases suggesting they were the first to crystallise. Olivine grains only contained spinel inclusions suggesting they crystallised afterward. Clinopyroxene contained inclusions of spinel, olivine and plagioclase and showed replacement to hornblende. Hornblende contained inclusions of olivine, clinopyroxene and plagioclase, while some plagioclase grains also contained inclusions of hornblende. Although the olivines and spinels were demonstrably the first phases to crystallise there is therefore some ambiguity in the subsequent order of crystallisation of phases. It appears that a population of plagioclase grains (An<sub>96–79</sub>) crystallised before clinopyroxene. After subsequent hornblende crystallisation, another population of plagioclase grains crystallised. The groundmass assemblage of olivine (Fo<sub>66–59</sub>) + clinopyroxene = plagioclase (An<sub>79–46</sub>) clearly formed outside the hornblende stability field. An obvious question is at what depth in the crust each population crystallised before eruption and whether crystals are primocrysts, antecrysts or xenocrysts.

The grain size arrangements, assemblages and textures of glomerocrysts are strikingly similar to the textures observed in cumulate crustal xenoliths brought to the surface by explosive eruptions in nearby Bequia ([Camejo-Harry et al., 2018](#)) and elsewhere in the Lesser Antilles ([Cooper et al., 2016](#); [Melekhova et al., 2017](#); [Stamper et al., 2014a](#); [Tollan et al., 2011](#)). This supports the proposal that glomerocrysts may be either extracts of cumulate residues or lower pressure accumulations of near-liquidus minerals from H<sub>2</sub>O saturated magmas that were incorporated following destabilisation of crystal mushes in the trans-crustal magmatic system ([Camejo-Harry et al., 2018](#)). The compositional similarity between glomerocrysts and phenocrysts suggests that a large proportion of phenocrysts may form by glomerocryst disaggregation. The large grain sizes (>2 mm) of various textural types of clinopyroxene, plagioclase and hornblende are indicative of a complex history. Lower degrees of undercooling would facilitate the slow growth of larger sized phenocrysts. Entrainment from the lower crust and subsequent transport to upper storage regions, reheating and possibly magma mixing would result in both reverse and normal zoning, sieve textures, reaction rims and resorption as the crystal cargo attempted to re-equilibrate to changing pressure ( $\pm$  temperature) during magma ascent.

At least three populations of isolated plagioclase phenocrysts clearly distinguished by grain size and An content, can be related to early and late plagioclase crystallisation. Large sized plagioclase phenocrysts ( $\geq 2$  mm) containing zoning patterns of sizeable homogenous cores (An<sub>94–89</sub>) with thin rims suggests conditions of crystallisation under conditions of low undercooling up until just before eruption, upon which conditions changed abruptly causing the crystallisation of compositionally different rims. Core crystallisation is likely to have occurred in deeper storage regions under near isobaric conditions over long time scales. Decompression-triggered crystallisation is a comparatively rapid process that promotes the crystallisation of small-sized plagioclases ( $< 500$   $\mu$ m, An<sub>89–64</sub>). Such crystallisation would have occurred as magmas ascended to shallow levels within the crust ( $< 100$  MPa), which is supported by the MI trace ([Fig. 10b](#)) and major element ([Fig. 10c](#)) signatures. The associated drop in pressure causes water to



**Fig. 12.** (a) Volatile concentrations of CO<sub>2</sub> and H<sub>2</sub>O in Kick-'em-Jenny melt inclusions. Isobars calculated using the MagmaSat solubility model (Ghiorso and Gualda, 2015) with a median melt inclusion composition and 950 °C. Melt inclusion hosts denoted as follows: diamond, olivine; square, clinopyroxene; triangle, hornblende. Vapour isopleths inserted as grey lines showing degassing trends for: (i) the most primitive rock composition (KEJ010 obtained from Devine and Sigurdsson (1995)) modelled to have originally contained 6 wt% H<sub>2</sub>O (labeled) and 1.5 wt% CO<sub>2</sub> in the vapour and (ii) the melt inclusion with the highest measured H<sub>2</sub>O content modelled to have originally contained 7 wt% H<sub>2</sub>O (labeled) and 1.5 wt% CO<sub>2</sub> in the vapour. (b) Melt inclusion entrapment pressures for a given analysis.

exsolve providing the undercooling that drives crystallisation. Other plagioclase grains (ranging >500 µm and <2 mm) exhibited oscillatory zoning and abundant sieve textures possibly obtained from fluctuating conditions resulting from reheating of shallower magmas by hotter deeper magmas and decreasing water pressures as phenocrysts grew en route from mid crustal storage regions to shallow levels.

Hornblende is uncommon as a phenocryst phase in basic rocks because of its instability at low pressures in basaltic magmas (Rutherford and Hill, 1993) and the increased crystallinity that often accompanies hornblende crystallisation from hydrous basalts which results in increased magma viscosity and reduced eruptability (Barclay and Carmichael, 2004). Its prevalence in Kick-'em-Jenny lavas suggests that magma transit through the crust from deep magma storage regions kept pace with the slope of the hornblende-in curve, enabling it to remain saturated over a wide pressure range (Allen and Boettcher, 1983). Eruption would have had to be sufficiently rapid to prevent complete hornblende resorption. The calculated hornblende-plagioclase temperatures are in good agreement with the coexistence of these two minerals according to the water-undersaturated experiments of Holloway and Burnham (1972) (975 °C, 250 MPa). For water-saturated conditions, hornblende and plagioclase are found to coexist only at 1000 °C, 50 MPa (Barclay and Carmichael, 2004).

In summary, we propose on the basis of crystal textures and complex crystallisation sequences that the Kick-'em-Jenny and Kick-'em-Jack crystal cargos could have resulted from polybaric crystallisation of magma batches as they ascend through the crust and interact. In such a situation the different slopes of crystal saturation curves for a hydrous basalt and the exact P-T trajectory followed during ascent can give rise to very complex zonation patterns. In this regard the presence of early crystallising Fo-rich olivine and late crystallising groundmass olivine is striking.

#### 4.2. A case for cold, primitive, wet basalts

Kick-'em-Jenny has erupted some unusually MgO-rich magmas (≤13 wt%), raising the possibility that these are truly primitive, mantle-derived melts. Alternatively, the magnesian character of Kick-'em-Jenny magmas might simply correspond to large amounts of accumulated antecrystic or xenocrystic material. A number of chemical and petrographic criteria can be used to distinguish between these possibilities.

Xenocryst (or antecryst) addition would be expected to lead to correlations between xenocryst proportions and bulk rock chemistry. For example, if mafic phenocrysts are xenocrystic, we would expect the most magnesian lavas to have the highest modal proportions. This is

observed for hornblende phenocrysts, leading previous authors to conclude that hornblende assimilation is an important process occurring beneath Kick-'em-Jenny (Devine and Sigurdsson, 1995). However, the extent of hornblende addition observed cannot adequately account for the major element chemistry of Kick-'em-Jenny basalts as the hornblende phenocrysts in basalts contain only 14.7 wt% MgO, on average (Supplementary Table S1e), a value only slightly greater than that of the basalts themselves. Moreover, the low Ni and Cr contents of hornblendes cannot drive basalts to the observed levels that are, instead, consistent with values for mantle-derived magmas. Addition of xenocrystic hornblende should lead to a reduction in incompatible trace element concentrations, e.g. for Zr, Rb and Ba, which do not occur in hornblende to a significant degree. Passmore et al. (2012) used similar arguments to assess xenocryst incorporation in porphyritic Icelandic lavas. At Kick-'em-Jenny we see no negative correlation between modal hornblende and incompatible trace element concentrations (Supplementary Fig. S3).

Finally, xenocryst incorporation would be expected to yield discrete overgrowth or resorption textures around xenocryst cores. Although some olivines have magnesian cores with abrupt overgrowths (Fig. 4c), most large phenocrysts of hornblende and clinopyroxene display complex zonation patterns, consistent with primocrysts that are attempting to keep pace with changing pressure and temperature during magma ascent. This is likely to be especially important for hornblende due to its upper temperature stability limit of ≤1100 °C in basaltic magmas (Fig. 4, Melekhova et al., 2015). Magmas that ascend at or close to this temperature will tend to graze the hornblende liquidus leading to abrupt changes in stability and modal proportion, resulting in complex zonation patterns. These lines of evidence lead us to the conclusion that the magnesian character of Kick-'em-Jenny lavas is not simply a product of hornblende (or other xenocryst) assimilation.

If Kick-'em-Jenny lava chemistry is not controlled by crystal accumulation, then it must reflect either differentiation of a parental primitive magma or represent a range of primitive compositions, i.e. a range of magmas consistent with derivation from a peridotitic mantle source. The presence of rare Fo<sub>90</sub> olivine cores is consistent with such a scenario. We can calculate the Fo content of olivine in equilibrium with Kick-'em-Jenny magnesian lavas using  $Kd_{FeTotal-Mg} = [Fe^{Total}/Mg]_{ol}/[Fe^{Total}/Mg]_{melt}$ . The presence of significant ferric iron needs also to be taken into account. Assuming that Fe<sup>3+</sup> does not enter the olivine structure, the apparent  $Kd_{FeTotal-Mg}$  will be reduced from the canonical value of 0.31 in proportion to the amount of ferric iron, i.e.  $Kd_{FeTotal-Mg} = 0.31 * (1 - Fe^{3+}/Fe^{Total})$ . As noted above, for the eleven Kick-'em-Jenny lavas of Sigurdsson and



Shepherd (1974) for which  $\text{Fe}_2\text{O}_3$  and FeO were analysed separately, calculated  $K_{\text{FeTotal-Mg}}$  values are in the range 0.21 to 0.24 and the calculated equilibrium olivine compositions range from  $\text{Fo}_{91}$  to  $\text{Fo}_{78}$ . For typical mantle with primary (unreacted)  $\text{Fo}_{92-89}$  olivine, as recorded by mantle xenoliths from nearby Grenada (Parkinson et al., 2003), these calculations indicate that all Kick-'em-Jenny lavas with  $\geq 8$  wt% MgO are plausibly primitive, i.e. in equilibrium with mantle peridotite. On this basis, we suggest that all Kick-'em-Jenny lavas more magnesian than this represent primitive magmas. Their chemical variation can be ascribed to differing melting conditions and source mineralogy, for example, in terms of the amount of hornblende initially present at the onset of melting. Conversely, magmas with  $< 8$  wt% MgO reflect distinct chemical differentiation trends that can be ascribed to crystallisation of the observed phenocryst assemblages, notably hornblende which serves to drive down both CaO and MgO. The continued increase in  $\text{Al}_2\text{O}_3$  below 8 wt% MgO (Fig. 8) indicates that the proportion of fractionating plagioclase must be minor, despite its relative abundance as a phenocryst.

The relatively low calculated phenocryst and eruption temperatures ( $1000 \pm 30$  °C) reported above, coupled with high  $\text{H}_2\text{O}$  contents ( $\leq 5$  wt %), of Kick-'em-Jenny lavas provide clues as to their mantle origins. For magmas that ascend rapidly through the crust, as suggested by the preservation of the hornblende phenocrysts, magma source and eruption temperatures are expected to differ only by the combination of adiabatic cooling ( $\sim 50$  °C/GPa; Carmichael (2002)) and release of latent heat of crystallisation ( $\sim 2.5$  °C per % crystallised; Blundy et al. (2006)). As these effects broadly cancel out in the case of Kick-'em-Jenny, we infer that the source melting temperatures for primitive basalts ( $> 8$  wt % MgO) was of the order 1000 °C. This is strikingly similar to the dehydration melting solidus of hornblende-peridotite. According to the experimental data of Niida and Green (1999), the volatile-absent, hornblende-lherzolite solidus ranges from 980 °C at 0.8 GPa to 1000 °C at 1.5 GPa. Under these conditions, near-solidus melts would be  $\text{H}_2\text{O}$ -rich and magnesian, although, to date, no experimental studies of hornblende-peridotite melting have reported glass compositions. It is noteworthy that the melting interval over which hornblende (pargasite) and melt coexist during melting is around 30 °C. Over this interval melt composition is expected to change significantly because of the incongruent nature of pargasite melting (Niida and Green, 1999). Thus chemical variation in primitive melts may result from isobaric melting in the presence of hornblende beneath Kick-'em-Jenny. The curious kink in whole rock CaO contents around 8–10 wt% MgO (Fig. 8) may reflect such source controls on basalt chemistry.

Melting of mantle peridotite with various modal proportions of hornblende can provide a plausible interpretation of Kick-'em-Jenny magnesian basalts. However, phenocryst temperature calculations for basalts (1972) are somewhat lower (964–1008 °C, Table 4 & Fig. 11) than groundmass (eruption) temperatures (1006–1023 °C). We attribute this difference to the effects of adiabatic cooling and latent heat of crystallisation, whose combined effect is very little change in temperature from source to surface. The exact P-T trajectory that ascending basalts will follow depends on the relative rates of ascent and crystallisation. At Kick-'em-Jenny we suggest that adiabatic cooling preceded significant crystallisation, giving rise to phenocryst temperatures that are slightly lower than eruption temperatures.

A primary origin for magnesian Kick-'em-Jenny basalts is consistent with the presence of primitive basalts on Grenada to the south and St. Vincent (Bouvier et al., 2008; Heath et al., 1998) to the north. However, the nature of those primitive basalts and the melting processes that generated them are very different from those inferred for Kick-'em-Jenny. For example primary basalts on St. Vincent are thought to derive from partial melting of peridotite at 1130 to 1180 °C in equilibrium with a hornblende-free residuum (Heath et al., 1998) at pressures of  $\sim 1.4$  GPa according to the multiple saturation experiments of Pichavant et al. (2002) and Melekhova et al. (2015). These conditions are more typical

of primitive arc basalts worldwide with eruption temperatures  $\sim 1160$  to 1200 °C (Tatsumi and Eggins, 1995). The mantle source depth of primitive Grenada basalts is less clear, although their rare earth element chemistry (Shimizu and Arculus, 1975) and high eruption temperatures (ol + cpx + spl one atmosphere liquidus = 1325 °C; Stamper et al. (2014a)) argue for a deep, garnet-bearing peridotite source, again markedly different for the source conditions of Kick-'em-Jenny hornblende basalts. The geographical proximity of Kick-'em-Jenny to Grenada (8 km apart) makes the difference in mantle source regions intriguing, perhaps reflecting the local structure and mineralogy of the sub-arc mantle and slab.

#### 4.3. Fate of dissolved volatiles

High concentrations of  $\text{H}_2\text{O}$  were measured in Kick-'em-Jenny magmas (up to 5.0 wt%), consistent with the striking abundance of hornblende in erupted lavas together with delayed plagioclase crystallisation. High  $\text{Al}_2\text{O}_3$  contents of hornblendes also suggest crystallisation at high  $\text{H}_2\text{O}$  pressure (Grove et al., 2003; Melekhova et al., 2015; Nandedkar et al., 2014; Pichavant and Macdonald, 2007; Sisson and Grove, 1993). Fig. 12a confirms that  $\text{H}_2\text{O}$ -rich conditions correspond to greater pressures (and hence depths) within the crust than  $\text{H}_2\text{O}$ -poor conditions, a consequence of decompression-driven degassing, rather than isobaric, volatile-saturated crystallisation. Trace element geochemistry of MIs clearly demonstrates the proclivity of plagioclase crystallisation for lower pressure,  $\text{H}_2\text{O}$  poor conditions (Fig. 10b). Strontium, as a proxy for plagioclase crystallisation, shows high concentrations in the melt at high  $\text{H}_2\text{O}$  contents where plagioclase crystallisation is hindered, and a decrease in Sr concentrations as  $\text{H}_2\text{O}$  contents decrease enough for plagioclase to crystallise. Such plagioclase crystallisation would have been facilitated by decompression, accompanied by  $\text{H}_2\text{O}$  loss (Fig. 10c) and may also account for the increase in incompatible elements with decreasing  $\text{H}_2\text{O}$  (Fig. 10a). Of the remaining trace elements measured, only yttrium is significantly compatible in hornblende and likely to be taken up within its crystal structure during crystallisation at higher  $\text{H}_2\text{O}$  contents. As water concentrations decrease, hornblende becomes unstable and begins to break down, and elements such as Y are returned to the melt, plausibly accounting for its enrichment at low  $\text{H}_2\text{O}$  contents (Fig. 10a). The fact that many MIs preserve decompression crystallisation signals suggests that these MIs maintain some connection with the matrix glass after incorporation in the host mineral, as previously suggested by Blundy and Cashman (2005) for silicic MIs from Mount St. Helens.

Although magmas lose their dissolved volatiles as a natural consequence of ascent and eruption (Metrich and Wallace, 2008), knowledge of the initial water content of magmas is important since it can affect various aspects of magmatic evolution from source to eruption. MIs trapped in the earliest forming crystals stand the best chance of preserving original  $\text{H}_2\text{O}$  concentrations (Plank et al., 2013; Wade et al., 2008; Zimmer et al., 2010). Such crystals would usually be olivine, but at Kick-'em-Jenny the MI with the highest measured water content (5 wt%) was hosted in clinopyroxene, while olivine-hosted MIs contain  $< 3$  wt%  $\text{H}_2\text{O}$  (Fig. 12a). There are no MIs preserved in the  $\text{Fo}_{90}$  olivines thought to represent the deepest crystallisation conditions. The observed  $\text{H}_2\text{O}$  range is appreciably lower than would be expected for dehydration melts of hornblende-peridotite, estimated to be on the order 10–12 wt% by Niida and Green (1999). If Kick-'em-Jenny basalts were indeed derived by hornblende-peridotite melting then the actual water (and potentially  $\text{CO}_2$ ) content would be significantly higher than 5 wt%. In that case, either MIs have lost their original  $\text{H}_2\text{O}$  upon stalling in upper crustal reservoirs causing inclusions to reset to lower  $\text{H}_2\text{O}$  concentrations (Buchholz et al., 2013) or their host minerals crystallised at lower pressures during magma ascent. Blundy et al. (2010) have shown that decompression crystallisation can lead to MIs trapping after a significant fraction of the original magmatic volatile content has been exsolved for the case where the initial dissolved  $\text{CO}_2$  content was relatively high. In this case, both the  $\text{H}_2\text{O}$  and  $\text{CO}_2$  contents

of Kick-'em-Jenny primitive magmas may have been considerably higher than those recorded by MI trapped during decompression crystallisation of olivine and clinopyroxene within the crust. In that case, the high CO<sub>2</sub> contents of some MIs might be an indication of the original magmatic contents, albeit their low H<sub>2</sub>O contents must have been significantly modified by diffusive processes, e.g. Bucholz et al. (2013).

The generally low CO<sub>2</sub> contents of Kick-'em-Jenny MIs points towards significant degassing on magma ascent. The degassing paths taken by MIs with the highest volatile contents from high to low pressures within the crust (Fig. 12a) implies that the magmatic system is highly pressurised with respect to H<sub>2</sub>O and CO<sub>2</sub> at depth. Using the MagmaSat model of Ghiorsio and Gualda (2015), we can model the path a primitive melt from depth would have followed during depressurisation in a closed system. Using the most magnesian bulk rock composition for Kick-'em-Jenny lavas (1972 KEJ010, Devine and Sigurdsson (1995)), we can vary the initial H<sub>2</sub>O and CO<sub>2</sub> content until the modelled degassing paths match our data. The best fit for our data using this composition requires magma to have been initially H<sub>2</sub>O rich (6 wt%) and CO<sub>2</sub>-rich (1.5 wt%), which is significantly higher than estimated from MIs (Fig. 12a), but still not as high as the primitive melt H<sub>2</sub>O contents for dehydration melting of hornblende-peridotite proposed by Niida and Green (1999). A similar trend calculated through the MI with the highest measured H<sub>2</sub>O content requires initially 7 wt% H<sub>2</sub>O and 1.5 wt% CO<sub>2</sub> (Fig. 12a). CO<sub>2</sub>-rich (2.5–3.2 wt%) MIs perhaps provide a rare glimpse of the deeper magmatic system (Fig. 12). Clearly further experimental studies of the major element composition and volatile contents of melts of hornblende-peridotite are required.

The presence of a volatile-charged system at depth beneath Kick-'em-Jenny has ramifications for hazard assessment especially if such volatiles are not periodically released and allowed to build up in the system. The fact that such high volatile contents were present in volcanic rocks (lavas) suggests that there must have been some mechanism for efficient outgassing to promote effusive behaviour. Dondin et al. (2017) propose that brittle deformation of the rocks making up Kick-'em-Jenny's edifice increases rock permeability allowing outgassing from magma reservoirs through fracture networks. Effusive behaviour will therefore be encouraged in lieu of explosive activity. These authors go on to suggest that explosive activity may be triggered by the presence of a low permeability lava dome capping the crater (which has been observed at this volcano in the past (Lindsay and Shepherd, 2005)) that can prevent the release of volatiles from the magmatic system. If the rate of magma ascent exceeds the rate of volatile release through fracture networks, explosive behaviour may also be favoured (Dondin et al., 2017).

#### 4.4. Changes in Kick-'em-Jenny between 1972 & 2014: implications for hazards

Because sample collection for 1972 and 2013/14 expeditions occurred some years after the last preceding documented eruption (Table 1), we cannot say with absolute certainty that crater lavas are exclusively representative of those eruptions. Sigurdsson and Shepherd (1974) describe the 1972 collection as fresh and lacking associated organic material signifying the occurrence of eruptions between 1966 and 1972. The same may have occurred between the 2001-documented eruption and the 2013/14 expeditions. We therefore treat these sample collection dates as minimum eruption periods.

1972 and 2013/14 crater and cone lavas contain similar mineral assemblages, textures and chemistries. However, there are certain differences between the eruptive products from each time cohort that indicate an evolution in the volcanic system over 40+ years. Firstly, the erupted rock types have changed from predominantly basaltic in 1972 (Devine and Sigurdsson, 1995) to basaltic andesite in composition in 2013/14 (Supplementary Fig. S1). MI compositions also show a broader range in 2013/14 deposits from basaltic andesite to dacite

while 1972 MIs span basalt to basaltic andesite. Thirdly, there is evidence for a reduction in the crystallisation temperatures of olivine. The Ca content of olivine has been shown to be primarily dependent on temperature in the presence of clinopyroxene (Shejwalkar and Coogan, 2013). The calculated isotherms using the algorithm of Shejwalkar and Coogan (2013) for orthopyroxene-bearing 2013/14 landslide deposits show that olivine crystallisation for Kick-'em-Jenny occurred over a temperature range of 900–1050 °C (Fig. 5b). As expected, compositional factors exercise some control on crystallisation temperatures, but what is more apparent is that some 2013/14 olivine crystallised under lower temperatures than 1972.

Omitting unusually high CO<sub>2</sub>-low H<sub>2</sub>O MIs, it is apparent that eruptions making up the 1972 sample collection derive from both deep and shallow magma storage regions, whereas the eruptions contributing to the 2013/14 samples tapped mainly the shallower reservoirs (Fig. 12b). This may have ramifications for flank stability at Kick-'em-Jenny. If recent eruptive activity is indeed concentrated at a shallower depth, the likelihood for a magmatic component triggering flank collapse is heightened. Even though an explosive eruption is unlikely to trigger a tsunami at the volcano's current elevation (Lindsay et al., 2005), shallow magma injection as a destabilising mechanism combined with the noted instability of the volcano's edifice (Allen et al., 2018; Dondin et al., 2017) increases its susceptibility for slope collapse which may trigger a tsunami.

We see a situation where the volcano's future eruptive compositions will likely become predominantly more silicic than previous eruptions as time progresses. The tendency for explosive or dome-forming eruptions, that may eventually lead to explosive eruptions as the permeability of the edifice is decreased, is also greater because of the accompanying increase in magma viscosity. If a significant episode of cone building does occur as a result of such compositional changes, the likelihood of a tsunami threat will be increased (Allen et al., 2018). The added complication of a shallow magma reservoir makes this volcano a prime candidate for multiple direct and indirect hazards to nearby populations on Grenada and the Grenadines.

#### 4.5. Linking geochemical variations to Kick-'em-Jenny's structural history

All mineral phase (except clinopyroxene) and bulk rock compositions from Kick-'em-Jenny volcano display consistent differences between crater/cone deposits (irrespective of sampling period) and landslide deposits. Although Kick-'em-Jack samples are too few to adequately compare bulk rock compositions, hornblende compositions (Fig. 6) display a remarkable similarity to Kick-'em-Jenny landslide deposits. These variations agree with current understanding of the different incarnations of Kick-'em-Jenny volcano. The constraints on sample collection for 2013/14 crater deposits and their similarity to 1972 deposits suggest that these samples characterize the current edifice of Kick-'em-Jenny. The locations of 2013/14 landslide deposits (possibly including Kick-'em-Jack landslide deposit samples), combined with their geochemical similarities, imply that these samples may be older than crater deposits and may be in fact more related to the volcano's proto edifice.

Before delving into geochemical distinctions, the first key difference lies in mineral assemblages (Fig. 2): 1972 and 2013/14 crater/cone deposits contain assemblages of olivine, clinopyroxene, hornblende, plagioclase and spinel whereas 2013/14 landslide deposits also contain these phases but with minor olivine and the addition of orthopyroxene. More evolved Kick-'em-Jack landslide deposits are also different, containing only hornblende, plagioclase and spinel. Kick-'em-Jenny 1972 and 2013/14 crater deposits display a separate trend of higher CaO contents for a given MgO from 2013/14 landslide deposits (Fig. 8b). This is indicative of different pressures of crystallisation as modelled for the M- and C-series lavas of Grenada by Stamper et al. (2014b). Lower pressures delay the onset of clinopyroxene saturation relative to olivine (Stamper et al., 2014b); initial, protracted crystallisation of olivine

enriches the melt in CaO. The result is a differentiation path matching 1972 and 2013/14 crater deposits, although, these do not reach CaO contents as high as Grenada C-series. Higher pressures allow near simultaneous crystallisation of clinopyroxene and olivine, depleting the melt in both CaO and MgO and producing a trend matching 2013/14 landslide deposits, similar to that seen in Grenada's M-series, albeit starting at lower initial MgO.  $\text{Al}_2\text{O}_3$  contents suggest that  $\text{H}_2\text{O}$  contents were higher, and plagioclase crystallisation was later, in melts producing crater deposits compared to landslide deposits (Fig. 8c). The absence of orthopyroxene in crater deposits may also be related to higher magmatic  $\text{H}_2\text{O}$  contents, which tends to stabilise olivine by reducing silica activity (Sisson and Grove, 1993). Hornblende-plagioclase crystallisation temperatures are comparatively lower in landslide deposits than crater deposits (Table 4).

We propose that 2013/14 landslide deposits were generated from melts originating at higher pressures than 1972 and 2013/14 crater deposits. Plagioclase crystallisation occurred later in crater than landslide deposits, possibly aided by higher  $\text{H}_2\text{O}$  contents. If landslide deposits are indeed representative of Kick-'em-Jenny's proto edifice, there appears to have been an upward migration in the depth of melt generation with this new incarnation of the current edifice. In many cases, Kick-'em-Jack landslide deposits match Kick-'em-Jenny landslide deposits, however, the limited bulk rock data are equivocal, showing similarity to the crater deposit trend or lying in between two trends (Figs. 8 & 9). It is very likely that Kick-'em-Jack landslide deposits are related to Kick-'em-Jenny deposits, but represent more evolved samples from an older Kick-'em-Jack edifice (plausibly proto Kick-'em-Jenny).

#### 4.6. Kick-'em-Jenny plumbing system

Measurements of dissolved volatiles in MIs provide only a minimum estimate of magma storage pressures if the melts were not volatile-saturated at the point of entrapment (Blundy and Cashman, 2008) or if significant degassing preceded MI trapping (Blundy et al., 2010). Pressures (Fig. 12b) can be converted to depths using an appropriate density model ( $2750 \text{ kg/m}^3$  for  $\geq 250 \text{ MPa}$  and  $2500 \text{ kg/m}^3$  for  $\text{MI} < 250 \text{ MPa}$ ). On the basis of MI volatile saturation pressures, we propose that Kick-'em-Jenny's crustal magmatic system consists of a main reservoir somewhere between 7 and 12 km depth (Fig. 13), with shallower networks of smaller melt lenses and crystal mushes interconnected by sub-vertical feeding structures. Rare  $\text{CO}_2$ -rich MIs contain low  $\text{H}_2\text{O}$  contents that suggest the interaction between stalled magma and ascending  $\text{CO}_2$ -rich vapours originating deep within the plumbing system. Evidence that this magmatic system extends throughout the crust comes also from hypocenters of pre-eruptive earthquakes collected from the 2015 and 2017 eruptions (obtained from UWI-SRC) that cover a similar depth range (Fig. 13).

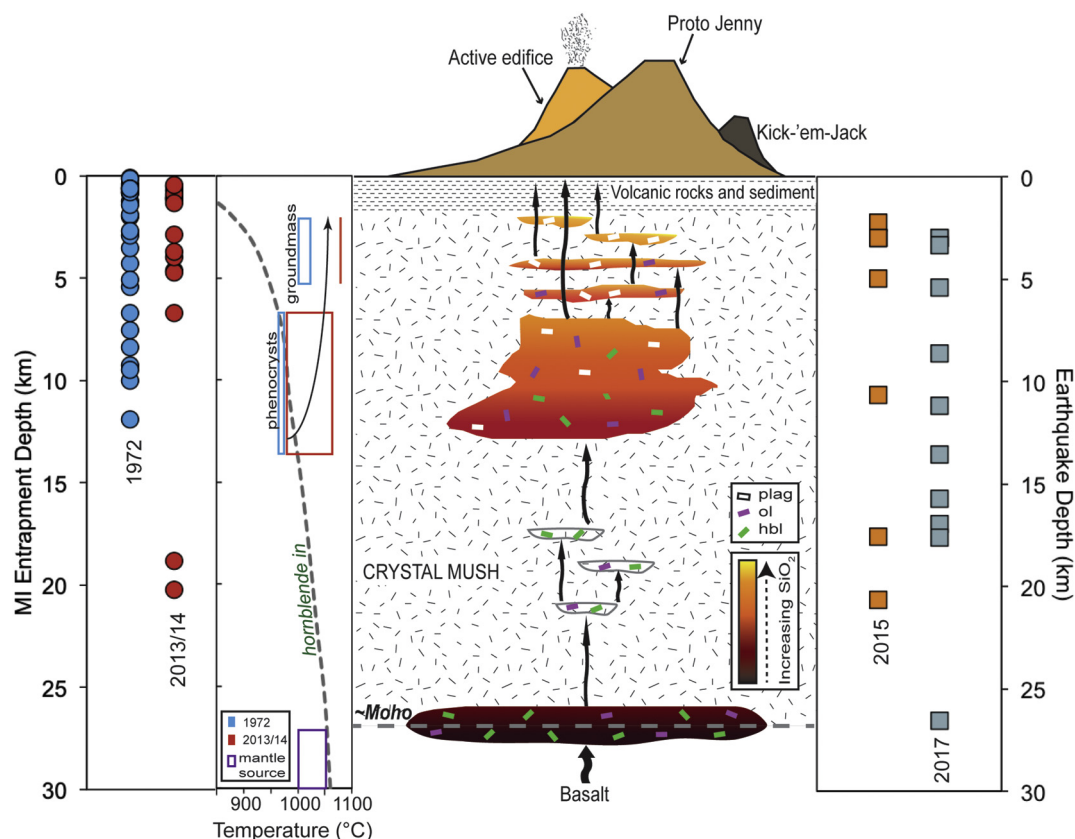
As for other arc systems, mantle-derived basalts are injected into the lower crust, upon which fractionation occurs forming cumulate piles and residual melts. For Kick-'em-Jenny, the temperature of melts coming out of the mantle, on the basis of the phenocryst and eruption thermometry results presented above, are unusually low ( $\sim 1000^\circ\text{C}$ ) given the Mg-rich nature of the least evolved basalts. Hydrous mantle melts can be generated from hornblende-peridotite sources at these temperatures via dehydration melting at pressures of around 1 GPa. These low temperatures promote hornblende stability in primitive basaltic melts and their derivatives, in contrast to much hotter primitive magmas from St. Vincent and Grenada. Because mantle melting involves incongruent breakdown of hornblende (Niida and Green, 1999), melts of hornblende-peridotite are likely to be enriched in hornblende components. In fact, as noted above, the bulk composition of the most primitive Kick-'em-Jenny lavas are quite similar to their hornblende phenocrysts. Because melts produced from hornblende-peridotite will have a high dissolved hornblende content they are likely to crystallise copious quantities of hornblende during ascent, provided that temperatures during

ascent lie within the hornblende stability field. There are no phase equilibrium studies of Kick-'em-Jenny basalts, but studies of a tholeiitic basalt in equilibrium with a mixed  $\text{H}_2\text{O}$ - $\text{CO}_2$  fluid ( $\text{XH}_2\text{O} = 0.59$ ) over a wide pressure range (Holloway and Burnham, 1972) show that the hornblende-in curve has a slope of  $\sim 125^\circ\text{C/GPa}$  in the pressure range 0.8 (1050  $^\circ\text{C}$ ) to 0.2 (975  $^\circ\text{C}$ ) GPa; below 0.2 GPa the hornblende liquidus curves sharply to much lower temperatures, reflecting its instability at low pressure. We suspect that the  $\text{XH}_2\text{O}$  in the source region of Kick-'em-Jenny primitive basalts was somewhat greater than that of 0.59, leading to slightly lower hornblende liquidus temperatures. For example, for the  $\text{H}_2\text{O}$ -saturated case, hornblende saturation temperatures are around 40  $^\circ\text{C}$  lower, but the corresponding slope on the hornblende-in curve is  $\sim 133^\circ\text{C/GPa}$ , i.e. 0.8 GPa, 1010  $^\circ\text{C}$  and 0.2 GPa and 930  $^\circ\text{C}$  (Krawczynski et al., 2012). Thus, near adiabatic ascent of primitive Kick-'em-Jenny basalts is likely to track the hornblende liquidus, producing hornblende crystals over a very broad pressure interval (Fig. 13). Olivine of contrasting Fo content will bracket hornblende crystallisation, from  $\text{Fo}_{90}$  close to the mantle source region to  $\text{Fo}_{70-60}$  in lower pressure microlites and crystal rims, following significant hornblende and clinopyroxene crystallisation. We suggest that this accounts for large size of Kick-'em-Jenny and Kick-'em-Jack hornblende phenocrysts and their complex zoning, as well as providing constraints on the ascent path of these magmas.

Primitive magmas ascend to feed middle and upper crustal reservoirs and in doing so become increasingly chemically evolved. This process may involve classic crystal fractionation, whereby melts and crystal become separated from each other, closed system crystallisation, or entrainment of crystal cargos from the surrounding lithologies. Crystal cargos are plausibly dominated by mafic phases at depth that are progressively replaced by felsic phases with decreasing vertical depth. Upper crustal reservoirs are likely sustained by the destabilisation of ponded magma in the mid-crust (Cashman et al., 2017) which may be initiated by diverse melts percolating the mush effectively driving fluctuating conditions of the system and maintaining instability. Magmas with  $>8 \text{ wt\% MgO}$  are considered to be diverse primitive melts from hornblende-bearing peridotite sources in the sub-Kick-'em-Jenny mantle. The diversity in melts with  $<8 \text{ wt\% MgO}$  (Table 3) may be due to fractional crystallisation, but might equally represent a range of primitive melts with different evolutionary histories and water contents that are allowed to percolate crystal mushes beneath Kick-'em-Jenny. Reversely zoned phases suggest the infiltration of hotter melts forming high temperature rims within magma reservoirs already containing more evolved melts. These varied compositions encourage the mixing of magma batches (for which we have seen evidence, Fig. 3f) allowing cumulates, glomerocrysts and antecrysts to be entrained and transported to different storage regions and eventually erupted in lava flows. The amount of volatiles contained will be determined by their initial abundance in mantle-derived basalts, which may become enriched upon magma differentiation and evolution. Magma ascent and later decompression in lower storage regions will cause volatile exsolution thereby increasing the bubble content of the magma.

The high initial volatile contents and highly pressurised, vertically extensive system beneath Kick-'em-Jenny make it plausible that volcanic eruptions may be driven by deeply generated volatiles, as proposed for Soufrière Hills, Montserrat, by Christopher et al. (2015). Increased magma pressure together with the mixing of diverse melts can lead to a major destabilisation of the trans-crustal magmatic system, which in turn can provide the overpressure required for an eruption. Periods of dormancy between eruptions are marked by a return of the system to a period of relative stability where the major fluxes are caused by basalt injection at depth which slowly evolves by the fractionation of segregated melts. As melt layers begin to connect and move upward in the system, more silicic magma are disrupted allowing volatiles to be liberated and magma pressure to be increased, returning the system to a period of unrest (Christopher et al., 2015).





**Fig. 13.** Schematic illustration of the sub-volcanic system beneath Kick-'em-Jenny volcano. The system is composed of a chief magma storage region (~7–12 km) and shallower reservoirs all connected by sub-vertical feeding structures. Shown on the left are entrapment depths for 1972 and 2013/14 phenocryst hosted melt inclusions. Density models used: 2750 kg/m<sup>3</sup> for melt inclusions trapped at  $\geq 250$  MPa and 2500 kg/m<sup>3</sup> for melt inclusions trapped at  $< 250$  MPa. The adjacent panel displays calculated olivine-liquid and hornblende-plagioclase thermometry for groundmass and phenocrysts respectively, temperature range for mantle source region of basalts and the inferred hornblende-in curve after Holloway and Burnham (1972). The arrow denotes inferred magma ascent P-T trajectory. Shown on the right are pre-eruptive earthquake depths for 2015 and 2017 eruptions. Errors on these depths are  $\pm 10$  km. Pre-eruptive events occurred July 12–22, 2015 and July 8–28, 2017 respectively (monitoring data provided by the UWI-SRC database). The Moho is placed at 27 km after Christeson et al. (2008). Mantle-derived hydrous basalts are injected at depth upon which fractionation and migration of residual melts to upper storage regions occurs. Melt compositions can be varied as denoted by SiO<sub>2</sub> content. At depths greater, melt lenses are inferred by grey polygons. With decreasing depths, crystal cargos and residual melts become progressively evolved.

## 5. Conclusion

Kick-'em-Jenny volcano has erupted basalts, basaltic andesites and andesites with glomeroporphyritic textures. Phenocryst phases differ according to the location of sample collection: crater deposits contain spinel, olivine, clinopyroxene, hornblende and plagioclase while landslide deposits contain similar phases with comparatively less olivine and the addition of orthopyroxene. Evidence for disequilibria such as zoning, breakdown rims and sieve and replacement textures are recurring features of phenocrysts. The H<sub>2</sub>O content of Kick-'em-Jenny magmas, as measured from melt inclusions (MIs), is  $\leq 5$  wt%, but may be significantly higher if MIs were trapped after significant ascent and degassing. Crystallisation temperatures, from a variety of thermometers, are surprisingly low for such MgO-rich basalts (~1000 °C). The high H<sub>2</sub>O concentration and low temperature explains the abundance of hornblende phenocrysts as well as high Al<sub>2</sub>O<sub>3</sub> contents and corresponding delayed plagioclase crystallisation. The presence of Fo<sub>90</sub> olivine phenocryst cores, MgO- and Ni-rich whole rock chemistry, and low eruption temperatures of Kick-'em-Jenny basalts is consistent with a mantle source of hornblende-peridotite that undergoes dehydration melting.

A comparison of 1972 and 2013/14 crater samples shows an evolution in erupted deposits from predominantly basaltic to basaltic andesite in composition. MI compositions have also expanded in range from basaltic to andesitic to basaltic andesite to dacitic in more recent eruptions. MI data provide evidence for a highly pressurised, vertically

extensive magma system originating in the mantle wedge but with a principal magma storage region ~7–12 km. This has implications for the type of eruptive activity expected especially if gases are not periodically exsolved from the magmatic system. The fact that such high volatile contents were measured in extrusive rocks suggests there is currently some mechanism for efficient outgassing to promote effusive behaviour. It is likely that this temporal compositional evolution of Kick-'em-Jenny magmas will have a control on the nature of future eruptions.

The spatial differences in sample collection have also revealed distinctions in rock and mineral geochemistry. We have used this to show that crater deposits from 1972 and 2013/14 sample Kick-'em-Jenny's current edifice, while 2013/14 landslide deposits likely represent proto Kick-'em-Jenny (an edifice existing pre 43.5 kyr BP). 2013 Kick-'em-Jack landslide deposits are also likely related to proto Kick-'em-Jenny substantiating previous suggestions that these volcanoes may belong to the same volcanic complex. The magmatic systems beneath these two incarnations of Kick-'em-Jenny have also changed over time, with the depth of melt generation decreasing and melt H<sub>2</sub>O content increasing between these periods.

Supplementary data to this article can be found online at <https://doi.org/10.1016/j.jvolgeores.2019.01.023>.

## Data access statement

All underlying data are provided in full within this paper, either in the main text or as accompanying supplementary material.

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